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# Unified Theory Calculations of Stark Broadened Hydrogen Lines Including Lower State Interactions 

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C. R. Vidal, J. Cooper, ${ }^{* *}$ and E. W. Smith

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#### Abstract

Recently published calculations of hydrogen Stark broadening on the basis of the unified classical path theory have been extended to include lower state interactions in the final line profile. A detailed comparison with experiments in the density range $10^{13-10^{17} \mathrm{~cm}^{-3}}$ is given.


Key words: Classical path; hydrogen lines; line wings; Stark broadening; unified theory.

## 1. Introduction

In a recent paper [1], ${ }^{1}$ henceforth referred to as paper I, the unified classical path theory [2] was generalized for the case of upper and lower state interactions and was applied to the Stark broadening of hydrogen. The thermal average of the time development operator and the final intensity profile of any hydrogen line were derived for the general case including lower state interactions. Numerical calculations for the thermal average including lower state interactions have been presented. The calculations of the final line profiles, however, have so far been restricted to the Lyman lines. In this paper we discuss calculations of the intensity profile including lower state interactions which are more involved because they require the evaluation of tetradic operators and contain more extensive sumations over vector coupling coefficients. The influence of lower state interactions is demonstrated for the first few lines of the Balmer series and possible simplifications for the higher series members are pointed out. A detailed comparison with various experiments covering the electron density range from about $10^{13}-10^{17} \mathrm{~cm}^{-3}$ is given requiring for low electron densities a convolution of the first series members with the Doppler profile. The unified theory calculations are also compared with static calculations. The applied computer program is presented in the appendix.

## 2. Basic Relations

In this section we summarize briefly the basic relations, which have been derived in paper I and are used for the unified theory calculations presented in this paper.

With the static ion field approximation the shape $I(\omega)$ of a Stark broadened line is given by

$$
\begin{equation*}
I(\omega)=\int_{0}^{\infty} P\left(\mathscr{E}_{i}\right) I\left(\omega, \mathscr{E}_{i}\right) d \mathscr{E}_{i} \tag{1}
\end{equation*}
$$

where the normalized distribution function $P\left(\mathscr{C}_{i}\right)$ is the low frequency component of the fluctuating electric microfields. In this manner we regard the radiator as being an atom subjected to a static field $\mathscr{E}_{i}$ and perturbed by the electrons. In the unified classical path theory [2] the ion field dependent line shape $I\left(\omega, \mathscr{C}_{i}\right)$ is obtained from

$$
\begin{equation*}
I\left(\omega, \mathscr{E}_{i}\right)=\frac{1}{\pi} \sum \operatorname{Im}\left\{d \frac{1}{\Delta \omega_{\mathrm{op}}\left(\mathscr{E}_{i}\right)-\mathscr{L}\left[\Delta \omega_{\mathrm{op}}\left(\mathscr{E}_{i}\right)\right]} \mathrm{d}\right\} \tag{2}
\end{equation*}
$$

$\mathbf{d}$ is the dipole moment. The matrix elements of $\Delta \omega_{\text {op }}$ specify the distance $\Delta \omega$ from a particular Stark component shifted by the static field $\mathscr{E}_{i}$ and $\mathscr{L}$ is essentially the Fourier transform of the thermal

[^0]average. Within the classical path approximation and the impact approximation $\mathscr{L}\left(\Delta \omega_{\mathrm{op}}\right)$ is given by
\[

$$
\begin{equation*}
\mathscr{L}\left(\Delta \omega_{\mathrm{op}}\right)=-i \Delta \omega_{\mathrm{op}} \int_{0}^{\infty} \exp \left(+i t \Delta \omega_{\mathrm{op}}\right) \overline{\mathscr{F}}(1)(t) d t \Delta \omega_{\mathrm{op}} \tag{3}
\end{equation*}
$$

\]

where the thermal average

$$
\begin{equation*}
\overline{\mathscr{F}}(1)(t)=n_{e} \int d \mathbf{x}_{1} d \mathbf{v}_{1} W\left(\mathbf{v}_{1}\right)\left[\mathscr{U}_{1}\left(\mathbf{R}, \mathbf{x}_{1}, \mathbf{v}_{1}, t\right)-1\right] . \tag{4}
\end{equation*}
$$

$n_{e}$ denotes the electron density and $W$ the velocity distribution function. The tetradic time development operator $\mathscr{U}_{1}$ is defined by the time ordered expression

$$
\begin{equation*}
\mathscr{U}_{1}\left(\mathbb{R}, \mathbf{x}_{1}, \mathbf{v}_{1}, t\right)=\mathscr{O} \exp \left\{-\frac{i}{\hbar} \int_{0}^{t} \tilde{\mathscr{V}}_{1}\left(\mathbf{R}, \mathbf{x}_{1}, \mathbf{v}_{1}, t^{\prime}\right) d t^{\prime}\right\} \tag{5}
\end{equation*}
$$

with the binary interaction

$$
\begin{equation*}
\tilde{\mathscr{V}}_{1}\left(\mathbf{R}, \mathbf{x}_{1}, \mathbf{v}_{1}, t\right)=\exp \left\{i t \mathscr{H}_{0} / \hbar\right\} \mathscr{V}_{1}\left(\mathbf{R}, \mathbf{x}_{1}, \mathbf{v}_{1}, t\right) \exp \left\{-i t \mathscr{H}_{0} / \hbar\right\} . \tag{6}
\end{equation*}
$$

$\mathscr{V}_{1}$ is the Coulomb interaction potential between the radiator and a single perturbing electron. The Hamiltonian $\mathscr{H}_{0}$

$$
\begin{equation*}
\mathscr{H}_{0}=\mathscr{H}_{a}+e Z \mathscr{C}_{i} \tag{7}
\end{equation*}
$$

consists of the Hamiltonian $\mathscr{H}_{a}$ of the unperturbed radiator and the static ion part $e Z_{\mathscr{C}}{ }_{i}$. (For details see eqs (IV.16)-(IV.18),(III.14)-(III.16) and (II.2) of paper I.)

In evaluating the preceding equations for well isolated hydrogen lines we have to consider matrix elements only between states with the same principal quantum number $n$ (no-quenching approximation) and it then is most convenient to work with parabolic states $|n q m\rangle$. $m$ is the magnetic quantum number and the quantum number $q$ is defined to be

$$
\begin{equation*}
q=n_{1}-n_{2} \tag{8}
\end{equation*}
$$

with $n_{1}$ and $n_{2}$ being the usual parabolic quantum numbers which have to satisfy the relation

$$
\begin{equation*}
n=n_{1}+n_{2}+|m|+1 . \tag{9}
\end{equation*}
$$

Furthermore, we distinguish quantum numbers which refer to the lower state from the upper state quantum numbers by a prime.

In taking matrix elements of the operators $\Delta \omega_{\mathrm{op}}, \mathbf{d}$ and $\mathscr{L}\left(\Delta \omega_{\mathrm{op}}\right)$ between parabolic states it was shown in paper I that for the general case of upper and lower state interactions we have the following relations. For $\Delta \omega_{\text {op }}$, which is diagonal in parabolic states, one obtains

$$
\begin{equation*}
\left\langle n^{\prime} q^{\prime} m^{\prime} ; n q m\right| \Delta \omega_{\mathrm{op}}\left|n^{\prime} q^{\prime} m^{\prime} ; n q m\right\rangle=\Delta \omega-\Delta \omega_{i}\left(n, q, n^{\prime}, q^{\prime}\right) \beta \tag{10}
\end{equation*}
$$

where $\Delta \omega$ is the frequency perturbation from the position of the unperturbed line, $\beta$ the normalized field strength in units of the Holtsmark field strength $\mathscr{E}_{0}$

$$
\begin{equation*}
\beta=\mathscr{E}_{i} \mid \mathscr{C}_{0} ; \mathscr{E}_{0}=\left(\frac{4 \pi}{3}\right)^{2 / 3} e n_{e}^{2 / 3} \tag{11}
\end{equation*}
$$

and

$$
\begin{equation*}
\Delta \omega_{i}\left(n, q, n^{\prime}, q^{\prime}\right)=\left(\frac{4 \pi}{3}\right)^{2 / 3} \frac{3}{2}\left(n q-n^{\prime} q^{\prime}\right) \frac{\hbar}{m} n_{e}^{2 / 3} \tag{12}
\end{equation*}
$$

$\Delta \omega_{i}$ is the frequency shift of a particular Stark component characterized by the quantum numbers $n, q, n^{\prime}$ and $q^{\prime}$ due to the Holtsmark field strength $\mathscr{E}_{0}$. The matrix elements of $\mathscr{L}\left(\Delta \omega_{\text {op }}\right)$ are given by

$$
\begin{align*}
& \left\langle n^{\prime} q_{b}^{\prime} m_{b}^{\prime} ; n q_{b} m_{b}\right| \mathscr{L}\left(\Delta \omega_{\mathrm{op}}\right)\left|n^{\prime} q_{a}^{\prime} m_{a}^{\prime} ; n q_{a} m_{a}\right\rangle=-i \pi\left[\Delta \omega-\Delta \omega_{i}\left(n, q_{b}, n^{\prime}, q_{b}^{\prime}\right) \beta\right]^{2} \\
& \sum_{q_{c} q_{c}^{\prime}}\left\langle n^{\prime} q_{b}^{\prime} m_{b}^{\prime} ; n q_{b} m_{b}\right| K\left(q_{c}, q_{c}^{\prime}\right)\left|n^{\prime} q_{a}^{\prime} m_{a}^{\prime} ; n q_{a} m_{a}\right\rangle i\left(\Delta \omega, \beta, n, n^{\prime}, q_{b}, q_{b}^{\prime}, q_{c}, q_{c}^{\prime}\right) \tag{13}
\end{align*}
$$

where

$$
\begin{align*}
& \left\langle n^{\prime} q_{b}^{\prime} m_{b}^{\prime} ; n q_{b} m_{b}\right| K\left(q_{c}, q_{c}^{\prime}\right)\left|n^{\prime} q_{a}^{\prime} m_{a}^{\prime} ; n q_{a} m_{a}\right\rangle=(-1)^{m_{a}+m b} \sum_{\substack{l_{a}^{\prime} l_{a}^{\prime} l_{b}^{\prime} L m_{c} m_{c}^{\prime}}}(2 L+1) \\
& \left\langle n^{\prime} q_{a}^{\prime} m_{a}^{\prime} \mid n^{\prime} l_{a}^{\prime} m_{a}^{\prime}\right\rangle\left\langle n^{\prime} l_{a}^{\prime} m_{c}^{\prime} \mid n^{\prime} q_{c}^{\prime} m_{c}^{\prime}\right\rangle\left\langle n^{\prime} q_{c}^{\prime} m_{c}^{\prime} \mid n^{\prime} l_{b}^{\prime} m_{c}^{\prime}\right\rangle\left\langle n^{\prime} l_{b}^{\prime} m_{b}^{\prime} \mid n^{\prime} q_{b}^{\prime} m_{b}^{\prime}\right\rangle \\
& \left\langle n q_{a} m_{a} \mid n l_{a} m_{a}\right\rangle\left\langle n l_{a} m_{c} \mid n q_{c} m_{c}\right\rangle\left\langle n q_{c} m_{c} \mid n l_{b} m_{c}\right\rangle\left\langle n l_{b} m_{b} \mid n q_{b} m_{b}\right\rangle \\
&  \tag{14}\\
& \binom{l_{a}^{\prime} l_{a} L}{-m_{c}^{\prime} m_{c} M^{\prime}}\binom{l_{a}^{\prime} l_{a} L}{-m_{a}^{\prime} m_{a} M}\binom{l_{b}^{\prime} l_{b} L}{-m_{c}^{\prime} m_{c} M^{\prime}}\binom{l_{b}^{\prime} l_{b} L}{-m_{b}^{\prime} m_{b} M} .
\end{align*}
$$

The unitary transformation $\langle n l m \mid n q m\rangle$ from parabolic to spherical states can be expressed in terms of $3 j$-symbols [3]

$$
\langle n l m \mid n q m\rangle=(-1)^{1 / 2(1+m-q-n)} \quad \sqrt{2 l+1}\left(\begin{array}{l}
\frac{n-1}{2} \frac{n-1}{2}  \tag{15}\\
l \\
\frac{m-q}{2} \frac{m+q}{2}-m
\end{array}\right)
$$

Introducing the normalized time

$$
\begin{equation*}
s=\tilde{\omega}_{p} t \tag{16}
\end{equation*}
$$

and the normalized frequency

$$
\begin{equation*}
\Delta \omega_{R}=\left[\Delta \omega-\Delta \omega_{i}\left(n, q_{b}, n^{\prime}, q_{b}^{\prime}\right) \beta\right] / \tilde{\omega}_{p} \tag{17}
\end{equation*}
$$

with $\tilde{\omega}_{p}=\sqrt{8 \pi n_{e} e^{2} / m}$ being the plasma frequency, the Fourier transform of the thermal average is defined to be

$$
\begin{equation*}
i\left(\Delta \omega_{R}, \beta, n, n^{\prime}, q_{b}, q_{b}^{\prime}, q_{c}, q_{c}^{\prime}\right)=\lim _{\epsilon \rightarrow 0} \frac{1}{\pi} \int_{0}^{\infty} \exp \left[\left(-\epsilon+i \Delta \omega_{R}\right) s\right] \bar{F}\left(s, n, q_{c}, n^{\prime}, q_{c}^{\prime}\right) d s \tag{18}
\end{equation*}
$$

The thermal average $\bar{F}(s)$ was evaluated in paper I for the general case of upper and lower state interactions and it was shown that $\bar{F}(s)$ may be approximated by a function $G(s)=\Sigma_{k} G_{k}(s)$ whose Fourier transform can be given analytically such that

$$
\begin{equation*}
i\left(\Delta \omega_{R}\right)=\sum_{k} i\left(k, \Delta \omega_{R}\right) \tag{19}
\end{equation*}
$$

For most practical situations it turns out to be sufficient to consider only the first term in the series

$$
\begin{equation*}
i\left(k=1, \Delta \omega_{R}\right)=a_{1} b_{1}^{2} e^{-i Z_{1}}\left\{i H_{0}^{(1)}\left(Z_{1}\right)+H_{1}^{(1)}\left(Z_{1}\right)\left[1-\frac{i}{2 Z_{1}}\right]\right\} \tag{20}
\end{equation*}
$$

where $H_{0}^{(1)}$ and $H_{1}^{(1)}$ are Hankel functions and

$$
\begin{equation*}
Z_{1}=b_{1} \Delta \omega_{R} \tag{21}
\end{equation*}
$$

because the higher order terms usually affect the final line profile by not more than 2 percent around $\Delta \omega \simeq \tilde{\omega}_{p}$. The constants $a_{1}$ and $b_{1}$ are given by

$$
\begin{equation*}
a_{1}=-4 \sqrt{\pi} n_{e} D^{3} C^{2}\left[B-\ln \left(4 C^{2}\right)\right] \tag{22}
\end{equation*}
$$

and

$$
\begin{equation*}
b_{1}=\frac{9}{4 \pi^{2}} C\left[B-\ln \left(4 C^{2}\right)\right]^{2} \tag{23}
\end{equation*}
$$

where $D$ is the Debye length, $B$ a constant of the order of unity (see table II in the appendix of paper I) and

$$
\begin{equation*}
C=\frac{3}{2}\left(n q-n^{\prime} q^{\prime}\right) \frac{\hbar}{m D v_{a v}} \tag{24}
\end{equation*}
$$

Finally we give for completeness the matrix elements of $d \otimes d$

$$
\begin{aligned}
& \left\langle n^{\prime} q_{a}^{\prime} m_{a}^{\prime} ; n^{\prime} q_{b}^{\prime} m_{b}^{\prime}\right| d \otimes d\left|n q_{a} m_{a} ; n q_{b} m_{b}\right\rangle=(-1) m_{a}^{\prime} \sum_{l_{a} l_{a}^{\prime}}\left\langle n^{\prime} q_{a}^{\prime} m_{a}^{\prime} \mid n^{\prime} l_{a}^{\prime} m_{a}^{\prime}\right\rangle\left\langle n l_{a} m_{a} \mid n q_{a} m_{a}\right\rangle \\
& \left(\begin{array}{ccc}
l_{a}^{\prime} & l_{a} & 1 \\
m_{a}^{\prime}-m_{a} \mu
\end{array}\right)\binom{l_{a}^{\prime} l_{a} 1}{0}\left[\left(2 l_{a}+1\right)\left(2 l_{a}^{\prime}+1\right)\right]^{1 / 2}\left\langle n l_{a}\right| r\left|n^{\prime} l_{a}^{\prime}\right\rangle
\end{aligned}
$$

where the radial matrix elements are given in eq (63.2) of Bethe and Salpeter [4].
In case of no lower state interaction where the time development operator between lower states is replaced by a unit operator the matrix elements of $\mathscr{L}\left(\Delta \omega_{\mathrm{op}}\right)$ in eq (13) and (14) simplify significantly and we have

$$
\begin{align*}
&\left\langle n^{\prime} q^{\prime} m^{\prime} ; n q_{b} m\right| \mathscr{L}_{u}\left(\Delta \omega_{\mathrm{op}}\right)\left|n^{\prime} q^{\prime} m^{\prime} ; n q_{a} m\right\rangle=-i \pi\left[\Delta \omega-\Delta \omega_{i}\left(n, q_{b}, n^{\prime}, q^{\prime}\right) \beta\right]^{2} \\
& \sum_{q_{c}}\left\langle n q_{b} m\right| K_{u}\left(q_{c}\right)\left|n q_{a} m\right\rangle i\left(\Delta \omega, \beta, n, n^{\prime}, q_{b}, q^{\prime}, q_{c}\right) \tag{26}
\end{align*}
$$

where

$$
\begin{equation*}
\left\langle n q_{b} m\right| K_{u}\left(\dot{q}_{c}\right)\left|n q_{a} m\right\rangle=\sum_{l_{a}, m_{c}} \frac{1}{2 l_{a}+1}\left\langle n q_{a} m \mid n l_{a} m\right\rangle\left\langle n q_{b} m \mid n l_{a} m\right\rangle\left[\left\langle n l_{a} m_{c} \mid n q_{c} m_{c}\right\rangle\right]^{2 .} \tag{27}
\end{equation*}
$$

Equation (26) simplifies even further for the Lyman lines with $\left|n^{\prime} q^{\prime} m^{\prime}\right\rangle=|100\rangle$. For this case numerical calculations were given in paper I. '

## 3. Properties of the $\mathscr{L}$ and $\mathbb{K}$ Matrix

In eqs (13) and (26) we have split the $\mathscr{L}$-matrix into two parts. One part, the $K$-matrix, contains all the vector coupling coefficients or $3 j$-symbols and is independent of the plasma parameters, while the other part, the Fourier transform of the thermal average $i(\Delta \omega)$, contains all the broadening parameters. The $K$-matrix is completely specified by the upper and lower state principal quantum numbers $n$ and $n^{\prime}$ and needs to be calculated only once for every hydrogen line.

From eq (14) we realize first of all that the $K$-matrix is symmetric to the diagonal

$$
\begin{equation*}
\left\langle n^{\prime} q_{b}^{\prime} m_{b}^{\prime} ; n q_{b} m_{b}\right| K\left(q_{c}, q_{c}^{\prime}\right)\left|n^{\prime} q_{a}^{\prime} m_{a}^{\prime} ; n q_{a} m_{a}\right\rangle=\left\langle n^{\prime} q_{a}^{\prime} m_{a}^{\prime} ; n q_{a} m_{a}\right| K\left(q_{c}, q_{c}^{\prime}\right)\left|n^{\prime} q_{b}^{\prime} m_{b}^{\prime} ; n q_{b} m_{b}\right\rangle \tag{28}
\end{equation*}
$$

while the $\mathscr{L}$-matrix is not, due to the factor $\left[\Delta \omega-\Delta \omega_{i}\left(n, q_{b}, n^{\prime}, q_{b}^{\prime}\right) \beta\right]^{2}$. Next we see from the $3 j$-symbols that $K$ is diagonal in the quantum number $M$.

$$
\begin{equation*}
M=m_{a}^{\prime}-m_{a}=m_{b}^{\prime}-m_{b} \tag{29}
\end{equation*}
$$

Hence we may arrange the $K$-matrix and also the $\mathscr{L}$-matrix in such a manner that they are block diagonal in $M$, where $M=0, \pm 1, \pm 2, \ldots, n+n^{\prime}-2$. We also notice that matrix elements which differ only in the sign of $M$ are identical.

$$
\begin{align*}
\left\langle n^{\prime} q_{b}^{\prime}-m_{b}^{\prime} ; n q_{b}-m_{b}\right| K\left(q_{c}, q_{c}^{\prime}\right) \mid n^{\prime} q_{a}^{\prime}-m_{a}^{\prime} ; & \left.n q_{a}-m_{a}\right\rangle \\
& =\left\langle n^{\prime} q_{b}^{\prime} m_{b}^{\prime} ; n q_{b} m_{b}\right| K\left(q_{c}, q_{c}^{\prime}\right)\left|n^{\prime} q_{a}^{\prime} m_{a}^{\prime} ; n q_{a} m_{a}\right\rangle \tag{30}
\end{align*}
$$

Consequently, blocks which differ in the sign of $M$, can also be made identical. This greatly simplifies the matrix inversion required by eq (2). The problem is further simplified by the fact that also the $d \otimes d$-matrix can be made block diagonal in

$$
\begin{equation*}
\mu=m_{a}-m_{a}^{\prime}=m_{b}-m_{b}^{\prime}, \tag{31}
\end{equation*}
$$

where $\mu$ can only take on the values 0 and $\pm 1$. As a result we finally have to evaluate only the blocks with $M=0$ and $M=1$ of the $K$ and $\mathscr{L}$-matrix, respectively, because the block with $M=-1$ gives identically the same contribution to the final line profile as the block with $M=1$ and all the other blocks with $|M|>1$ do not contribute due to the $d \otimes d$ matrix. Further useful symmetry properties of the $K$-matrix are given in the following equations:

$$
\begin{align*}
&\left\langle n^{\prime} q_{b}^{\prime} m_{b}^{\prime} ; n q_{b} m_{b}\right| K\left(q_{c}, q_{c}^{\prime}\right)\left|n^{\prime} q_{a}^{\prime} m_{a}^{\prime} ; n q_{a} m_{a}\right\rangle \\
&=\left\langle n^{\prime}-q_{b}^{\prime}-m_{b}^{\prime} ; n-q_{b}-m_{b}\right| K\left(q_{c}, q_{c}^{\prime}\right)\left|n^{\prime}-q_{a}^{\prime}-m_{a}^{\prime} ; n-q_{a}-m_{a}\right\rangle  \tag{32}\\
&=\left\langle n^{\prime} q_{b}^{\prime} m_{b}^{\prime} ; n q_{b} m_{b}\right| K\left(-q_{c},-q_{c}^{\prime}\right)\left|n^{\prime} q_{a}^{\prime} m_{a}^{\prime} ; n q_{a} m_{a}\right\rangle
\end{align*}
$$

The latter relation simplifies our summation over $q_{c}$ and $q_{c}^{\prime}$ in eq (13) because also for the Fourier transform $i(\Delta \omega)$ we have

$$
\begin{equation*}
i\left(\Delta \omega, \beta, n, n^{\prime}, q_{b}, q_{b}^{\prime}, q_{c}, q_{c}^{\prime}\right)=i\left(\Delta \omega, \beta, n, n^{\prime}, q_{b}, q_{b}^{\prime},-q_{c},-q_{c}^{\prime}\right) \tag{33}
\end{equation*}
$$

In calculating $I\left(\omega, \mathscr{E}_{i}\right)$ according to eq (2) we always have to invert complex matrices. Since for computational purposes it is more convenient to work with real matrices and since we need only the imaginary part of the inverted matrix we make use of the following relations
where

$$
\begin{align*}
& X+i Y=[A+i B]^{-1} \\
& X=\left[A+B A^{-1} B\right]^{-1}  \tag{34}\\
& Y=-\left[B+A B^{-1} A\right]^{-1}
\end{align*}
$$

with $X, Y, A$ and $B$ being real matrices. From eqs (2), (13) and (18) it is clear that $A$ contains the diagonal matrix of $\Delta \omega_{\mathrm{op}}$ and the sine transform of the thermal average and $B$ the cosine transform of the thermal average. In the limit of large $\Delta \omega$ we have

$$
\begin{equation*}
A=B+\Delta \omega \cdot I \tag{35}
\end{equation*}
$$

(see eq (X.18) of paper I) which simplifies to lowest order the $Y$-matrix to

$$
\begin{equation*}
Y \simeq-B / \Delta \omega^{2} \tag{36}
\end{equation*}
$$

This is the one electron limit, which as pointed out already in section IV. of paper I, does not require a matrix inversion.

## 4. The Influence of Lower State Interactions and the Static Limit

In the figures $1,2,3,7$ and 8 the intensity of the line is plotted versus the frequency perturbation $\Delta \omega$ in units of the plasma frequency $\tilde{\omega}_{p}=\sqrt{8 \pi n e^{2} / m}$. The frequency scale is preferred, because it represents essentially the energy perturbation and is therefore more meaningful for the discussion of Stark broadening than, for example, the wavelength scale, which is usually more convenient in the measurement. Furthermore, $\Delta \omega / \tilde{\omega}_{p} \leqq 1$ is essentially the domain of the unmodified impact theory.

We first demonstrate the influence of lower state interaction on the final line profile in figures 1 and 2 and compare three different methods of evaluation. In the first, most general method lower state interactions are taken into account and the line profiles are calculated on the basis of eqs (13) and (14) using the correct Stark effect for electrons and ions (solid curves in fig. 1). In the second method based on eqs (26) and (27) we neglect lower state interactions for the electrons, i.e., we do


Figure 1. The influence of lower state interactions on the Balmer lines $H_{\alpha}$ to $H_{\delta}$.


Figure 2. Comparison of the Balmer and Lyman lines with the upper state principal quantum numbers $\mathrm{n}=6$, 8 and 10.
not allow for perturbations of the lower state sublevels by electrons (dashed curves in fig. l, solid curves in fig. 2). However, we still use the correct Stark effect for the static ions. In the third and simplest method the line profiles are calculated like in case of the Lyman lines with no influence of the lower state for the electrons as well as the static ions (dashed curves in fig. 2). These Lyman profiles have been presented previously in Paper I.

We realize first of all from figure 1 that in view of the accuracy to be expected from the unified classical path theory we may neglect lower state interaction due to the electrons for $\mathrm{H}_{\delta}$ and therefore also for all the higher Balmer lines. For $n_{e}=1.3 \cdot 10^{13} \mathrm{~cm}^{-3}$ and $T=1850 \mathrm{~K}$ the biggest calculated difference between the cases with and without lower state interaction is 43 percent for $\mathrm{H}_{\alpha}$, 10 percent for $\mathrm{H}_{\beta}, 6$ percent for $\mathrm{H}_{\gamma}$ and 4 percent for $\mathrm{H}_{\delta \delta}$. These differences have been obtained for the distant, almost quasistatic wing. They become slightly smaller in the purely static wing and are significantly smaller in the line center ( $14 \%$ for $\mathrm{H}_{\alpha}, 9 \%$ for $\mathrm{H}_{\beta}$, and $3 \%$ for $\mathrm{H}_{\gamma}$ and $2 \%$ for $\mathrm{H}_{\delta}$ ). Similar differences may be obtained at any electron density and temperature: In the distant, purely static wing the difference will always be identical ( $39.2 \%$ for $\mathrm{H}_{\alpha}, 9.3 \%$ for $\mathrm{H}_{\beta}, 5.7 \%$ for $\mathrm{H} \gamma$ and $3.6 \%$ for $\mathrm{H}_{\delta}$ ), because the static ions are treated in both cases with the same, correct Stark effect, while the static electrons are treated in the second method with the Stark effect of the corresponding Lyman line.

In figure 2 the Balmer lines and the Lyman lines with the upper state principal quantum numbers $n=6,8$ and 10 are compared. In the line center we recognize the fact that the even Balmer lines have no unshifted Stark component, while the even Lyman lines do. The difference in the wings is also clear from the Stark effect of the Balmer and Lyman lines (see also table I of the paper by Vidal [5]). It therefore becomes apparent that for principal quantum numbers of about $n>10$ one may neglect the influence of the lower states on the final line profile altogether and use the Lyman profiles throughout.

Next we compare in figure 3 the results of the unified classical path theory (solid curves) with quasistatic calculations (dashed curves) in order to show to what extent quasistatic calculations may be useful. The even Balmer lines up to $n=14$ are plotted. The quasistatic profiles are based


Figure 3. Comparison of the unified theory and the quasistatic theory for the even Balmer lines with $\mathrm{n}=4$ ton $=14$.
on the low frequency component of the electric microfield distribution functions $[6,7]$ with a shielding parameter

$$
\begin{equation*}
r_{0} / D=0.0898 n_{e}^{1 / 6} / \sqrt{T} \tag{37}
\end{equation*}
$$

and the total density $N$ being $N=2 n_{e}$. In addition, short vertical lines mark the position of the average Weisskopf frequency

$$
\begin{equation*}
\Delta \omega_{c}=2 m v_{\mathrm{av}}^{2} /\left(3 \pi^{2} \bar{n}_{k} \hbar\right) \tag{38}
\end{equation*}
$$

as defined by Unsöld [8], where the average splitting

$$
\begin{equation*}
\bar{n}_{k}=\sum_{k} n_{k} f_{k} / \sum_{k} f_{k} \tag{39}
\end{equation*}
$$

with $f_{k}$ being the oscillator strength of the $k$ th Stark component is frequently approximated by $\bar{n}_{k} \simeq n(n-1) / 2$ (see table I of the paper by Edmonds, Schlüter and Wells [9]). The Weisskopf frequency indicates the domain of the quasistatic theory and there have been various estimates, which all agree within a factor of 2 to 3 . From figure 3 it is apparent that with increasing principal quantum numbers the profiles based on the unified classical path theory approach more and more the static profiles and that the Weisskopf frequency turns out to be a rather conservative estimate for the useful range of the quasistatic theory. In the example presented we recognize that except for the very line center one may use quasistatic calculations almost throughout the entire profile for principal quantum numbers of $n \geq 10$. For practical purposes the hole in the center of the static profile and the structure in the line center, which for the higher series members appears to oscillate around the profile based on the unified classical path theory, will be smeared out by a convolution with the Doppler profile.

## 5. Comparison With Experiments and Other Theories

We start our comparison of the unified theory with experiments in the high and low electron density range with recent measurements by Wiese, Kelleher, and Paquette [10] performed on a
high current, wall-stabilized arc. The experimental setup and the method of evaluation is similar to the one described by Wiese, Paquette, and Solarski [11]. However, a number of refinements are incorporated in the experimental setup and the achievable accuracy has been greatly improved.

In figure 4 and 5 the red and blue wings of the measured $\mathrm{H}_{\beta^{-}}$and $\mathrm{H}_{\gamma}$-profiles are plotted (solid lines), which have been obtained in the same run assuring the same electron density and temperature for both profiles. For the moment we do not consider the asymmetries of the profile, since the theory is not yet arranged to describe them. The electron density and temperature for this particular run have been determined in several ways from the absolute intensity of $H_{\beta}$ with

$$
n_{e}=7.5 \cdot 10^{16} \mathrm{~cm}^{-3} \quad T=12570 \mathrm{~K}
$$

from the absolute intensity of $\mathrm{H}_{\gamma}$ with

$$
n_{e}=7.7 \cdot 10^{16} \mathrm{~cm}^{-3} \quad T=12640 \mathrm{~K}
$$

and from the absolute intensity of the continum in the visible and in the UV, which differ by the contribution from the Balmer continuum, with

$$
n_{e}=8.0 \cdot 10^{16} \mathrm{~cm}^{-3} \quad T=12730 \mathrm{~K}
$$



Figure 4. Comparison between the unified theory for $H_{\beta}$ and the experimental profle as measured by Wiese, Kelleher and Paquette.


Figure 5. Comparison between the unified theory for $H_{\gamma}$ and the experimental profile as measured by Wiese, Kelleher and Paquette.

The three values of the electron density differ in all runs by similar amounts. In particular the difference between the values from the $\mathrm{H}_{\beta}$ and the continuum intensity is very reproducible and gives rise to a 6 percent difference in the electron density. There are indications (private communication of Dr. Wiese) that these differences may be due to small non LTE (local thermodynamic equilibrium) effects in the arc because of its rather small dimension. This would be consistent with the validity criteria for LTE in inhomogeneous stationary plasmas (see section 6-10 of Griem [12]). Since the continuum intensity is least affected, the electron density and temperature as obtained from the continuum have been adopted as the best values, in particular, since the same values have been obtained from the continuum in the visible and the UV. For these parameter values the profiles as calculated with the unified theory including lower state interactions have been plotted in figures 4 and 5 with the same normalization $\int_{-\infty}^{+\infty} I(\Delta \lambda) d \lambda=1$ as the experimental profiles. We realize that the calculated profiles are slightly higher and narrower than the measured profiles. In both cases, however, the best agreement between theory and experiment in a least squares sense has been obtained with $8.5 \cdot 10^{16} \mathrm{~cm}^{-3}$, meaning that the theoretical value appears to be 6 percent larger than the experimental value.

Applying recent calculations of Kepple and Griem [13], one obtains from the half-, quarter-, and eighth-width $n_{e}=7.38 \cdot 10^{16} \mathrm{~cm}^{-3}$ for $\mathrm{H}_{\beta}$ and $n_{e}=8.05 \cdot 10^{16} \mathrm{~cm}^{-3}$ for $\mathrm{H}_{y}$ (as evaluated by Dr. Wiese), which reveals an intrinsic inconsistency because the two electron densities differ by about 10 percent.

We also notice that the worst agreement is in the very line center, a feature which is so far common to all impact theories. This fact makes the determination of the electron density on the basis of fractional widths rather questionable because its definition may be ambiguous and it effectively normalizes wing intensities with respect to the intensity in the very line center, where the theory seems to be least reliable. Although this is the most convenient and most widely used method, it is definitely preferable to determine the electron density from a least squares fit of the experimental and theoretical profile using the same normalization for both profiles.

In figure 6 the calculations of Kepple and Griem [13] are compared with our calculations for $\mathrm{H}_{\beta}$ and $n_{e}=6.4 \cdot 10^{16} \mathrm{~cm}^{-3}$ and $T=12.200 \mathrm{~K}$ indicating that our profile is narrower and higher,


Figure 6. Comparison between the unified theory with $B$ as obtained for $\rho_{\text {min }}=\lambda+\frac{3}{2} \mathrm{n}^{2} \mathrm{a}_{0}$, with $B=1.27$ as used by Kepple and Griem and the modified impact theory of Kepple and Griem.
hence giving rise to larger electron densities. In the very line center the most important difference appears to be due to the constant $B$ in eqs (22) and (23). While our calculations are based on a quantum number dependent constant $B$ for a lower cutoff $\rho_{\text {min }}=A+\frac{3}{2} n^{2} a_{0}$, which has been selected according to the validity criterion of the classical path theory, Kepple and Griem use a
larger value $B=1.27$ for all the Stark components. We have repeated our calculations with their value of the constant $B$. The profile is included in figure 6 and does not quite agree with their calculations. It should be pointed out that the normalization of the profile of Kepple and Griem is slightly too small, which may account for part of the remaining difference in the very line center, where the unified classical path calculations should go over to the results of the impact theory, if the ion field dependent cutoff (see eq. VI. 4 of paper I) is neglected.

From the preceding comparison we realize that at this stage the most important problem seems to be to obtain better values of the constant $B$ whose dependence on various cutoff procedures was discussed in the appendix of paper I for classical path theories neglecting time ordering. We note that in case of the $\mathrm{H}_{\beta}$ and $\mathrm{H}_{\gamma}$ profiles in figures 4 and 5 , better agreement between theory and experiment may be obtained with a constant $B$ larger than that used in this paper, and that in case of the Lyman- $\alpha$ experiment of Boldt and W. S. Cooper [14] discussed in paper I, better agreement is obtained with a smaller constant $B$.

For the moment we postpone the discussion of the various possibilities, which may affect the constant $B$, until we also have compared our results with measurements of the Balmer lines $\mathrm{H}_{3}-\mathrm{H}_{14}$ and the Paschen lines $\mathrm{P}_{6}-\mathrm{P}_{13}$ performed by Vidal [5, 15]. Unfortunately, these experiments do not provide an independent electron density measurement, which is as accurate as the measured profiles. Since, however, the measurements revealed a $\Delta \omega^{-5 / 2}$-wing, which for most of the lines extended over two orders of magnitude in intensity, the electron density was determined assuming that these measured $\Delta \omega^{-5 / 2}$-wings were identical to the asymptotic Holtsmark wings. In this manner the same electron density of $n_{e}=1.3 \cdot 10^{13} \mathrm{~cm}^{-3}$ was obtained within $\pm 4$ percent for all the Balmer lines from $\mathrm{H}_{4}-\mathrm{H}_{14}$ (see Vidal [5]). In paper I it was pointed out (on the basis of the Lyman line calculations) that the unified theory calculations actually give the $\Delta \omega^{-5 / 2}$-wings which were measured in the experiment and which extend much further into the line center than a quasistatic theory would predict. However, over the intensity range measured, these $\Delta \omega^{-5 / 2}$-wings do not necessarily coincide with the asymptotic Holtsmark wing.

In the following we reevaluate the measured line profiles employing now the complete line profile from the line center to the wings. We concentrate primarily on the Balmer lines, which have been measured more accurately than the Paschen lines and where we also have more series members available. The measurements were performed on a stationary radio frequency discharge within a magnetic bottle as described by Schlüter [16]. The electron temperature, which within the error limits is equal to the ion temperature, was measured to be $T=1850 \mathrm{~K}$. Consequently, we have to consider first of all to what extent Doppler broadening and the Zeeman effect may influence the first series members. In figure 7 the theoretical line profiles are shown for $n_{e}=1.3 \cdot 10^{13} \mathrm{~cm}^{-3}$ and $T=1850 \mathrm{~K}$ before (dashed lines) and after (solid lines) the convolution with the Doppler profile. It shows that up to around $\mathrm{H}_{8}$ Doppler broadening has to be taken into account. In estimating the influence of the Zeeman effect we notice that for a magnetic field of 1800 gauss ( 0.18 tesla) as typically used in the experiments, the separation of the outer components of the Lorentz triplet amounts to $0.168 \mathrm{~cm}^{-1}$. Neglecting complications due to the combined Zeeman and Stark effect this indicates that from around $\mathrm{H}_{7}$ (full half width $=1.76 \mathrm{~cm}^{-1}$ ) we may neglect the Zeeman effect for all the higher series members.

A comparison of the experimental and theoretical profiles for the Balmer lines $\mathrm{H}_{7}-\mathrm{H}_{14}$ gives the best agreement for $n_{e}=1.15 \cdot 10^{13} \mathrm{~cm}^{-3}$, a value which is slightly smaller than the value of $n_{e}=1.3 \cdot 10^{13} \mathrm{~cm}^{-3}$ given by Vidal [5]. This is not surprising because within the measured intensity range the apparent $\Delta \omega^{-5 / 2}$-wings of the unified theory calculations are all lying above the asymptotic Holtsmark wings on which the electron density value of Vidal [5] was based. For this new value the measured and calculated Balmer lines $\mathrm{H}_{7}-\mathrm{H}_{14}$ all agree within 5 percent over the entire line profile. These maximum deviations are comparable to the achieved accuracies of the measurement and are so small that they hardly show up in a plot like in the figures $1,2,3,7$, and 8 . For the lower Balmer lines only the line wings can be evaluated since the line center is noticeably affected by the Zeeman effect not included in the calculations. The agreement is not quite as good. However, the differences do not exceed 10 percent for the wings of $\mathrm{H}_{\beta}$ to $\mathrm{H}_{\delta}$, which appears to be partially due to the growing influence of the apparatus profile for the narrower profiles of the lower series members.

In figure 8 our caleulations are compared with the calculations of Kepple and Griem [13], which have been extended to $\mathrm{H}_{12}$ by Bengtson, Kepple and Tannich [17]. We see that their profiles are wider and that the higher series members do not show the measured $\Delta \omega^{-5 / 2}$-wings. From a comparison with figure 3 we also realize that with increasing principal quantum number $n$ our profiles approach the static profiles while the profiles of Kepple and Griem do not. Their profiles actually appear to differ more and more from the static profiles with increasing principal quantum number $n$, which manifests the fact that the modified impact theory is not able to describe the static electrons.


Figure 7. The Balmer line profiles as calculated with the unified theory for $\mathrm{n}_{\mathrm{e}}=1.3 \cdot 10^{13} \mathrm{~cm}^{-3}$ and $T=1850 K$ with and without the convolution of the Doppler profile.


Figure 8. Comparison between the unified theory and the modified impact theory of Kepple and Griem for the even Balmer lines $H_{6}$ to $H_{12}$.

As pointed out already in paper I, we emphasize again that contrary to the $\mathrm{H}_{\beta^{-}}$and $\mathrm{H}_{\gamma}$-profiles discussed above, the higher series members are insensitive to the exact value of the constant $B$. The reason is that these profiles are predominantly static in nature. In other words it means that the average Stark effect splitting, which is approximately equal to the half width of the higher series members, is significantly larger than the constants $a_{1}$ for the different Stark components (see eq (22)), which determine essentially the electron impact broadening. Since the line centers of the first series members are predominantly Doppler broadened, these low density experiments give no detailed information on the constant $B$. Hence, as one would expect, we also obtain a better agreement between theory and experiment than we did for the $\mathrm{H}_{\beta^{-}}$and $\mathrm{H}_{\gamma}$-profiles discussed above, where the agreement could have been worse due to the uncertainty of the constant $B$. This fact is particularly important for astrophysical applications where one is mainly interested in low density profiles, because these profiles are insignificantly affected by the remaining uncertainty of the constant $B$. One may, therefore, calculate low density profiles with rather large confidence.

## 6. Discussion

Since the value of the constant $B$ appears to be the most vague and restricting quantity in the current classical path theories, we will now summarize the effects which may influence the constant $B$ and are not included in the present calculations. First of all, we can state in general that the exact value of the constant $B$ is least important if either the static broadening exceeds the electron impact broadening, as discussed above for the low density profiles, or if according to eq (22) the value of $\ln \left(4 C^{2}\right)$ is much larger than the uncertainty of the constant $B$ which means that

$$
\begin{equation*}
-\ln \left(4 C^{2}\right) \gg 1 \tag{40}
\end{equation*}
$$

According to eq (24) the latter situation is most likely to occur for the innermost Stark components of any hydrogen line.

The constant $B$ is sometimes misleadingly referred to as the strong collision parameter. This is only partially correct, because its value depends not only on the lower cut off parameter $\rho_{\text {min }}$
but also on the upper cut off parameter $\alpha D$, where $D$ is the Debye length and the constant $\alpha$ has been varied in the literature from 1.1 (Griem, Kolb and Shen [18]) down to 0.6 (Chappell, J. Cooper and E. Smith [19]). As discussed in the appendix of paper I, this parameter $\alpha$ will also influence the limits on the time integral $\int V\left(t^{\prime}\right) d t^{\prime}$ in the time development operator $\mathscr{U}$, which for the $S$-matrix limit are extended from $-\infty$ to $+\infty$. In the unified theory calculations these limits have been extended from $-T$ to $+T$, where

$$
\begin{equation*}
T=\sqrt{\alpha^{2} D^{2}-\rho^{2}} / v \tag{41}
\end{equation*}
$$

in order to make the time limits consistent with the upper cut off parameter $\alpha D$. We also saw from the appendix of paper I that, as a function of the parameter $\alpha$, the constant $B$ varies as

$$
\begin{equation*}
B=B(\alpha=1)+2 \ln \alpha . \tag{42}
\end{equation*}
$$

The correct value of the constant $\alpha$ has not yet been determined conclusively. Since for large impact parameters one is dealing only with weak collisions, it has to be possible to determine the parameter $\alpha$ from a second order classical path theory, which means essentially within the frame work of the ordinary impact theory.

The influence of small impact parameter collisions on the constant $B$ is a much more involved problem. First of all, it is clear that one has to worry about collisions with impact parameters

$$
\begin{equation*}
\rho \widetilde{<} \rho_{\min }=\lambda+\frac{3}{2} n^{2} a_{0}=\lambda\left(1+3.78 \cdot 10^{-3} n^{2} \sqrt{T}\right) \tag{43}
\end{equation*}
$$

where classical path theories start to break down. As discussed for example, by E. W. Smith, C. R. Vidal and J. Cooper [20] we have to distinguish in this range between "completed" strong collisions, which contribute primarily to the impact limit, and "incompleted" strong collisions, which mostly contribute to the static wings. While the latter collisions are properly treated in a classical path theory and have no influence on the value of the constant $B$, the "completed" strong collisions can only be treated correctly by a quantum mechanical calculation, which has not yet been done.

Besides the limitation imposed by a classical path approach, we also have to consider the influence of time ordering which has been neglected in our time development operator $\mathscr{U}$ of eq (5). We notice that time ordering will become important if $V \tau / \hbar \geq 1$. With a typical collision time $\tau \approx \rho / v$, this is the case for impact parameters

$$
\begin{equation*}
\rho \widetilde{<} \rho_{\text {ord }}=\frac{3}{2}\left(n q-n^{\prime} q^{\prime}\right) \star . \tag{44}
\end{equation*}
$$

For most practical cases, where the value of $B$ becomes critical, $\rho_{\text {ord }}$ is larger than $\rho_{\min }$. Hence, there will be a domain, in which time ordering is important even within the region of validity of a classical path theory.

In the $S$-matrix limit the effect of time ordering has been investigated within the classical path theory for $\mathrm{Ly}_{\alpha}$ (Bacon, Shen, and J. Cooper [21]) and for $\mathrm{H}_{\alpha}$ (Bacon [22]) by solving the complete set of coupled differential equations, which define all the matrix elements of the time development operator. These calculations have demonstrated that the effect of time ordering may change the value of the constant $B$ by as much as $\pm 1$ and that the effect on the final line profile may amount to a few percent in intensity. These calculations have also included higher multipole terms. If one examines the effect of the higher multipole terms with respect to the case, where one considers only the dipole terms, one finds out that in case of Lyman- $\alpha$ as well as in case of $\mathrm{H}_{\alpha}$ the higher multipole terms lower the value of all the different constants $B$. However, at the electron density and temperature of these experiments (which are essentially the same), the change in the constant $B$ due to the higher order multipole terms affects the final line profile only very little.

For the more general case of the time development operator $\mathscr{U}$, required by the unified theory, the effect of time ordering has recently been examined for Lyman $\alpha$ [23]. These calculations show the results of the $S$-matrix limit in the line center and the decreasing influence of time ordering with increasing frequency pertubation $\Delta \omega$.

Due to the various outlined reasons which may affect the value of the constant $B$, it is rather difficult to make a quantitative statement of the possible error. If we assume that for impact parameters $\rho \gtrsim \rho_{\text {ord }}$ all the collisions which contribute to the value of $B$, have been treated correctly, one may give an upper bound on the possible error in $B$ based on the unitarity condition of the time
development operator. Since the time development operator may only oscillate within the limits of $\pm 1$, the integral over impact parameters from 0 to $\rho_{\text {ord }}$, which is weighted by the impact parameter $\rho$, will change the constant $B$ at the most by an amount of the order of unity. In addition to this the influence of the upper cut off parameter according to eq (42) has to be kept in mind. We also recall that in order to obtain the best agreement between theory and experiment we have to change the constant $B$ for the Lyman $\alpha$ profile [14] by an amount which is slightly larger than 1, and for the $\mathrm{H}_{\beta}$ and $\mathrm{H}_{\gamma}$ profile by an amount, which is smaller than 1. It should also be noted that the ion field splitting is not fully taken into account in the time development operator as discussed in section VI of paper I. It was shown that this may lead to slight modifications in the line center but not in the line wings. A full treatment would be rather difficult because the ion field exponentials in eq (VI. 4) of paper I remove the spherical symmetry of the problem and a simple cut off procedure as suggested by Kepple and Griem [13] does not appear to be completely adequate since it affects the normalization of the line profile. In addition, near the line center we would expect the quasistatic description for the ions to be invalid. This however, is practically unimportant because the Weisskopf frequency for the ions, which gives the range of validity for a quasistatic approach, is for low densities well inside the Doppler width and at high densities well inside the half width. Notice also that the breakdown of the no-quenching approximation may cause a small effect on the strong collision term (see ref. [20], p. 415).

So far we have ignored profile asymmetries which have been observed, for example, in Ly- $\alpha$, $\mathrm{H}_{\beta}$ and $\mathrm{H}_{\gamma}$. In the following we give a brief discussion in which we do not consider the well known asymmetries due to the $\omega^{4}$-factor in the expression for the power spectrum, the frequency to wavelength conversion, and the Boltzmann factor [24], which was neglected by assuming the elements of the atomic density matrix to be constant for all initial states. Asymmetries due to these effects will grow with increasing frequency perturbation $\Delta \omega$ and are always negligible in the line center. Hence, we are mainly concerned with the higher order multipole terms and the higher order Stark effect terms due to the electrons and ions.

In section VII of paper I it was shown how the higher order multipole terms due to the electron perturbers can be included in the interaction potential of the time development operator and it was pointed out that within the no-quenching assumption only a finite number of multipole terms exist. The main aggravation is that the unitary transformation which diagonalizes the time development operator after the spherical average is no longer a simple rotation. In the impact limit this will affect primarily the constant $B$ as noted above and cause only negligible asymmetries like in case of the time ordered solutions [21, 22].

The main reason for asymmetries especially in the line center has been shown to be due to the ions and we refer to a recent paper of Sholin [25] which supersedes the earlier papers by Griem [26] and Nguyen-Hoe, Drawin and Herman [27]. In this paper, it has been successfully demonstrated that the observed asymmetries of the Ly- $\alpha, \mathrm{H}_{\beta}$ - and $\mathrm{H} \gamma$-lines can all be explained within a static ion approach, where the main effect is caused by the quadrupole term of the ions and to a smaller extent by the quadratic Stark effect and the ion field dependent transition probability corrections of the individual Stark components. The electrons were assumed to cause only collision broadening and to introduce no further asymmetries.

It should be pointed out that the results presented by Sholin can easily be incorporated into the calculations presented here and work on this is in process. Like Sholin, we regard the radiator as being an atom in a static ion field $\mathscr{E}_{i}$ and perturbed by the electrons. Hence, one only has to modify the matrix elements of $\Delta \omega_{\mathrm{op}}$ in eq (10) and of $d \otimes d$ in eq (25) according to the relations given by Sholin, in order to include the higher order multipole terms and the higher order Stark effect terms due to the static ions.

If we incorporate the higher multipole terms due to the electrons in the time development operator together with those due to the ions, another interesting feature of the unified theory is that a number of the higher order multipole terms due to the electrons and ions like, for example, the most important quadrupole terms, will cancel each other in the static limit of the electrons. This effect was already suspected by Griem [12] (p. 94). However, this sort of mutual compensation occurs only if electrons and ions are treated within the framework of the same approximations (for example, quasi-static or impact) as pointed out by Sholin [25]. For this reason the occurrence of asymmetries gives important information on the type of electron broadening. It is also clear why asymmetries have been observed in the high density profiles of $L y_{\alpha}, H_{\beta}$ and $H \gamma$, where the ions may be treated quasi-statically and the electrons are over most of the measured line profile in the domain of the impact theory, but not in the low density profiles of the higher Balmer and Paschen lines, where ions and electrons are both in the quasi-static domain over almost the entire line profile.

Finally, we may conclude that, at this stage, the accuracy of the unified theory calculations depends primarily on the extent to which the final line profile is affected by the constant $B$ and
to a lesser extent on the asymmetries not yet included. Both effects are in turn determined by the electron density and temperature. It is probably safe to say that in general the electron density obtained with the unified theory in its present form will differ at the most by 10 percent from its true value. In the impact limit significantly better agreements between theory and experiment, which have been reported in the literature, have to be regarded as fortuitous. However, better results are definitely obtained for the higher, still well isolated series members and/or at low electron densities ( $n_{e} \leqslant 10^{13} \mathrm{~cm}^{-3}$ ), which are of particular astrophysical interest.

## 7. References

[1] Vidal, C. R., Cooper, J., and Smith. E. W., J. Quant. Spectr. Radiative Transfer, 10, 1011 (1970).
[2] Smith, E. W., Cooper, J., and Vidal, C. R., Phys. Rev. 185, 140 (1969).
[3] Hughes, J. W. B., Proc. Phil. Soc. 91, 810 (1967).
[4] Bethe, H. A., and Salpeter, E. E., Quantum Merchanics of One- and Two-Electron Atoms (Springer Verlag, Berlin 1957).
[5] Vidal, C. R., Proc. of the 7th Intern. Conf. on Phenomena in Ionized Gases, p. 168 (1966).
[6] Baranger, M., and Mozer, B., Phys. Rev. 115,521 (1959): 118,626 (1960).
[7] Hooper, C. F., Phys. Rev. 165,215 (1968).
[8] Unsöld, A., Physik der Sternatmosphären (Springer Verlag, Berlin 1955).
[9] Edmonds, F. N., Schlüter, M., and Wells, D. C., Mem. Roy. Astr. Soc. 71, 271 (1967).
[10] Wiese, W. L., Kelleher, D. E., and Paquette, D. R., to be published. (The authors would like to thank Dr. Wiese for making his measurements accessible to us prior to publication).
[11] Wiese, W. L., Paquette, D. R., and Solarski, J. E., Phys. Rev. 129, 1225 (1963).
[12] Griem, H. R., Plasma Spectroscopy (McGraw-Hill Book Co., Inc., New York, 1964).
[13] Kepple, P., and Griem, H. R., Phys. Rev. 173,317 (1968). see also University of Maryland, Report 831.
[14] Boldt, G., and Cooper, W. S., Z. Naturforschung $19 \mathbf{a}, 968$ (1964).
[15] Vidal, C. R., Z. Naturforschung 19a, 947 (1964).
[16] Schlüter, H., Z. Naturforschung 15a, 281 (1960).
[17] Bengtson, R. D., Kepple, P., and Tannich, J. D., Phys. Rev. A1, 532 (1970).
[18] Griem, H. R., Kolb, A. C., and Shen, K. Y., Astrophys. J. 135, 272 (1962).
[19] Chappell, W. R., Cooper, J., and Smith, E. W., J. Quant. Spectr. Radiative Transfer 9, 149 (1969).
[20] Smith. E. W., Vidal, C. R., and Cooper, J., J. Res. Nat. Bur. Stand. (U.S.), 73 A (Phys. and Chem.), No. 4, 405 (1969).
[21] Bacon, M., Shen, K. Y., and Cooper, J., Phys. Rev. 188, 50 (1969).
[22] Bacon, M., to be published.
[23] Godfrey, T., Vidal, C. R., Cooper, J., and Smith, E. W., Phys. Rev., to be published.
[24] Huber, D. L., and VanVleck, J. H., Rev. Mod. Phys. 38, 187 (1966).
[25] Sholin, G. V., Opt. Spectrosc. 26, 275 (1969).
[26] Griem, H. R., Phys. Rev. 140, A1140 (1965).
[27] Nguyen-Hoe, Drawin, H. W., and Herman, L., J. Quant. Spectr. Radiative Transfer 4, 847 (1964).
[28] Vidal, C. R., Cooper, J., and Smith, E. W., Nat. Bur. Stand. (U.S.), Monogr. 116, 143 pages (May 1970).

## 8. Appendix. Program for Calculating the Line Profile $I(\Delta \omega)$

The program presented has been arranged in a very similar manner to the program given in Appendix C of the NBS Monograph [28], which was only able to calculate the Lyman lines. The first general version (program version A) is in principle able to compute any hydrogen line and is only limited by the available memory space of the computer. It has been written to calculate the Lyman lines up to $n=16$, the Balmer lines up to $n=8$ and the Paschen lines up to $n=5$. By replacing the main program STBRHY and the function AIIM by the program STBRHY and the function AIIM listed in the end, the program may be simplified to calculate any hydrogen line neglecting lower state interaction (program version B). The arrays have been dimensioned such that the program is able to calculate all the Lyman, Balmer, and Paschen lines up to $n=16$, As demonstrated in figure 1, this simplified version is completely adequate for all the Balmer lines beyond $\mathrm{H}_{\delta}$. For all the Lyman lines it is also noticeably faster than the first general version and the program given in Appendix C of the NBS Monograph.

1. The Fourier transform of the thermal average.

The function FOUTR calculates essentially the Fourier transform of the thermal average such that

$$
\begin{equation*}
\mathrm{CCREAL}=\operatorname{Re}\left\{i \Delta \omega_{R}^{2} i\left(\Delta \omega_{R}, \beta, n, n^{\prime}, q_{b}, q_{b}^{\prime}, q_{c}, q_{c}^{\prime}\right)\right\} \tag{Al}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathrm{CCIMAG}=\operatorname{Im}\left\{i \Delta \omega_{R}^{2} i\left(\Delta \omega_{R}, \beta, n, n^{\prime}, q_{b}, q_{b}^{\prime}, q_{c}, q_{c}^{\prime}\right)\right\} \tag{A2}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathrm{DOM}=\Delta \omega_{R}=\left[\Delta \omega-\Delta \omega_{i}\left(n, q_{b}, n^{\prime}, q_{b}^{\prime}\right) \beta\right] / \tilde{\omega}_{p} \tag{A3}
\end{equation*}
$$

In evaluating $i\left(\Delta \omega_{R}\right)$ according to eq (19) it computes $i\left(k=1, \Delta \omega_{R}\right)$ as defined in eq (20) and $i\left(k=2, \Delta \omega_{R}\right)$ as defined in eq (X.22) of paper I. The required Bessel functions $J_{0}, J_{1}, Y_{0}$ and $Y_{1}$ are obtained from the subroutine BSJY01. For large and small arguments $\Delta \omega_{n}$ the proper expansions are applied as discussed in paper I . The constants $p_{1}, p_{2}$ and $b_{1}(P 1, P 2, B 1)$ specifying $i\left(k=1, \Delta \omega_{k}\right)$ and $a_{2}$ and $b_{2}(A 2, B 2)$ specifying $i\left(k=2, \Delta \omega_{R}\right)$ are set in the function AIIM and are calculated once for all the Stark components in the main program STBRHY.

## 2. The profile $I(\Delta \omega, \beta)$.

The function AIIM evaluates $I(\Delta \omega, \beta)$ as required by eq (1). The program is executed twice to obtain the contribution to $I(\Delta \omega, \beta)$ from the blocks of the $K$ and $\mathscr{L}$-matrix with $M=0 \quad(M B L=1)$ and $M= \pm 1(M B L=2)$. It establishes first of all the $\mathscr{L}$-matrix according to eq (13). The necessary $K$-matrix (three dimensional array SLOP) as defined in eq (14) is read in or calculated once in the main program, while the Fourier transform $i\left(\Delta \omega_{R}\right)$ is obtained from the subroutine FOUTR described above. In order to reduce the size of the array SLOP the symmetry of the $K$-matrix with respect to the diagonal (see eq (28)) is used.

The matrix oraering is established once in the main program to allow for the partitioning discussed in section 3 . The quantum numbers $q_{a}, m_{a}, q_{a}^{\prime}$ and $m_{a}^{\prime}$ are specified by the running index NRA, $q_{b}, m_{b}, q_{b}^{\prime}$ and $m_{b}^{\prime}$ by NRB and $q_{c}$ and $q_{c}^{\prime}$ by NRC. Having set up the real and imaginary matrices of the operator [ $\Delta \omega_{\mathrm{op}}-\mathscr{L}\left(\Delta \omega_{\mathrm{op}}\right)$ ] the imaginary part of the invers is determined according to eq (34) performing the matrix inversion by means of the subroutine ENVERS. The resulting matrix is then multiplied by the dipole matrix of eq (25), which is also calculated once in the main program such that

$$
\begin{align*}
& \text { DIPOL }=(-1)^{\frac{2}{2}\left(n+n^{\prime}-q-q^{\prime}-M-2\right)} \quad(2 l+1)\left(2 l^{\prime}+1\right)\langle n l| r\left|n^{\prime} l^{\prime}\right\rangle \\
& \left(\begin{array}{ll}
\frac{n-1}{2} \frac{n-1}{2} & l \\
\frac{m-q}{2} \frac{m+q}{2}-m
\end{array}\right)\left(\begin{array}{lll}
\frac{n^{\prime}-1}{2} \frac{n^{\prime}-1}{2} & l^{\prime} \\
\frac{m^{\prime}-q^{\prime}}{2} \frac{m^{\prime}+q^{\prime}}{2} & -m^{\prime}
\end{array}\right)\left(\begin{array}{lll}
l^{\prime} & l & 1 \\
m^{\prime}-m-M
\end{array}\right)\left(\begin{array}{lll}
l^{\prime} & l & 1 \\
0 & 0 & 0
\end{array}\right) \tag{A4}
\end{align*}
$$

The function AIIM listed at the end for the case of no lower state interactions is arranged almost identically to the function AIIM just described. The only difference is that now the $\mathscr{L}$ - and $K$ matrices are block diagonal in $m, m^{\prime}$ and $q^{\prime}$ according to eq (26) which allows a partitioning into smaller submatrices whose contribution to $I(\Delta \omega, \beta)$ is then again evaluated successively.
3. The final profile $I(\Delta \omega)$.

The main program STBRHY calculates the final line profile by performing the ion field average according to eq (1). It reads in first the low frequency component of the electric microfields [6, 7] for $0<\beta \leqslant 30$ in steps of 0.1 for 5 different values of the shielding parameter $r_{0} / D=0 ., 0.2,0.4,0.6$ and 0.8. A listing of the distribution functions is given at the end of the program. (The deck of the distribution function was kindly supplied by Dr. C. F. Hooper.) The program then reads in the density $n_{e}$, the temperature $T$, the upper and lower state principal quantum numbers $n$ and $n^{\prime}$, the initial value $\Delta \omega$, for which the line intensity is calculated, the logarithmic step width, the total number of points to be calculated, a parameter, which specifies the number of ion field integration points and finally 6 numbers, which specify the characteristic constants $a_{2}$ and $b_{2}$ of the $G_{2}$-function for all Stark components. These 6 numbers may be obtained from the calculations of the thermal average described in appendix A of the NBS Monograph. As shown, however, in paper I the contribution of the $G_{2}$-function can in practically all cases by neglected and hence one may set all 6 numbers to zero as done in this paper.

The program starts out to calculate the microfield distribution function for the particular shielding parameter as defined in eq (37) by using the 5 point interpolation subroutine POLY5 and stores the values in the array FIELD. It then calculates the average ion field

$$
\begin{equation*}
\beta_{\mathrm{av}}=\int_{0}^{\infty} \beta W(\beta) d \beta \tag{A5}
\end{equation*}
$$

by means of Weddle's rule (subroutine WEDDLE).
Next the program reads in a characteristic number NPUNCH. If this number is larger than 1
the program reads in the complete $K$-matrix (array SLOP) for the particular hydrogen line, which may have been obtained from another run of the program. Some numbers, which specify the size of the $K$-matrix, are read together with NPUNCH. If NPUNCH is smaller than 1 the program calculates the required $K$-matrix and if NPUNCH is equal to 1 it also punches the $K$-matrix out on cards. In the latter two cases the numbers NUR(1), NUR(2), IPI and NNCCQ, which specify the size of the $K$-matrix, may be set to any arbitrary value. The described feature was built into the program because the computation of the $K$-matrix may be very time consuming and is required only once for every hydrogen line as explained in section 3. It took, for example, on the CDC 38003 minutes to calculate the $K$-matrix of $\mathrm{H}_{\alpha}, 8$ minutes for $\mathrm{H}_{\beta}, 40$ minutes for $\mathrm{H}_{\gamma}$ and 100 minutes for $\mathrm{H}_{\delta}$.

As a next step the program calculates the wavelength in standard air of the particular hydrogen line. It then calculates all the necessary radial matrix elements $\langle n l| r\left|n^{\prime} l^{\prime}\right\rangle$ by means of the function RADMAT, which in turn obtains the factorials from the function FCTRL. The values are stored in the array RDM according to

$$
\mathrm{RDM}=\left(\begin{array}{lll}
l & l^{\prime} & 1  \tag{A6}\\
0 & 0 & 0
\end{array}\right)\langle n l| r\left|n^{\prime} l^{\prime}\right\rangle
$$

and the total line strength

$$
\left.\mathrm{STOT}=\sum_{l, l^{\prime}}(2 l+1)\left(2 l^{\prime}+1\right)\left(\begin{array}{ccc}
l & l^{\prime} & 1  \tag{A7}\\
0 & 0 & 0
\end{array}\right)^{2}|\langle n l| r| n^{\prime} l^{\prime}\right\rangle\left.\right|^{2}
$$

is determined. The program then calculates all the necessary unitary transformations $\langle n l m \mid n q m\rangle$ as defined in eq (15) using the $3 j$ symbol function S3J and stores them for the upper and lower states in the arrays SJUU and SJLL. With these transformations the dipol matrix elements between parabolic states are calculated and stored in the array DIPOL according to eq (A4). Simultaneously, the matrix ordering of the $K$ - and $\mathscr{L}$-matrix is established such that the quantum numbers $q, q^{\prime}, m$ and $m^{\prime}$ of a particular state $\left|n q m ; n^{\prime} q^{\prime} m^{\prime}\right\rangle$ are replaced by one running index NRUN and the block number MBLOCK, which may be 1 or 2 for $M=0$ and $M= \pm 1$ (see eq (29)). The values of the individual quantum numbers are stored in the arrays NQUU, NQLL, MMUU and MMLL.

The program then proceeds to calculate the $K$-matrix (array SLOP) according to eq (14) if NPUNCH is smaller or equal to 1 as explained above. Although the values for the eight different unitary transformations in eq (14) are stored, the computer time may still become rather long for the higher series members due to the extensive sumation over the quantum numbers $l_{a}, l_{a}^{\prime}, l_{b}, l_{b}^{\prime}$, $L, m_{c}$ and $m_{c}^{\prime}$.

As a last preliminary calculation the constants $p_{1}, p_{2}, b_{1}, a_{2}$, and $b_{2}$ for all the Stark components as required by the function AIIM for the subroutine FOUTR are determined and stored in the array FPAR. The constants $p_{2}, a_{1}$, and $b_{1}$ are evaluated for a lower cutoff $\rho_{\min }=\lambda+\frac{3}{2} n^{2} a_{0}$. As explained in the appendix of paper I the constant $B$ in eqs (22) and (23) is obtained from

$$
\begin{equation*}
B=2-3 \gamma-2\left[\frac{1-\cos (z)}{z^{2}}+\frac{\sin (z)}{z}-C i(z)\right] \tag{A8}
\end{equation*}
$$

where

$$
\begin{equation*}
z=3 n_{k} \pi / \rho_{\min } \tag{A9}
\end{equation*}
$$

The necessary cosine integral is calculated by the function COSINT.
Finally the program is prepared to actually perform the ion field average. The distribution function $P(\beta)$ is obtained from the function $W F L D$, which uses a 5 point interpolation on the values initially stored in the array FIELD. For $\beta>30$ the function $W F L D$ uses an asymptotic expansion. The ion field integral is performed by means of Weddle's rule (subroutine WEDDLE). The integrand is subdivided into several intervals depending on the spacing and the width of the individual Stark components. A convenient change of variables is used in every interval.
For

$$
\begin{equation*}
\Delta \omega>5 \beta_{\mathrm{av}} \Delta \omega_{i}\left[\left(n q-n^{\prime} q^{\prime}\right)_{\max }\right] \tag{A10}
\end{equation*}
$$

the ion field integral is replaced by $I\left(\Delta \omega, \beta=\beta_{\mathrm{av}}\right)$ as discussed in paper I. Besides the ion field integral the program calculates the wavelength perturbation $\Delta \lambda$ (DLAM), which corresponds to $\Delta \omega / \widetilde{\omega}_{p}$, the static profile for $N=n_{e}$ (FHOLTS), the asymptotic $\Delta \omega^{-5 / 2}$-wing (WING52), the one electron limit for $\beta=0$ (AWING), $I(\Delta \omega, \beta=0)$ and $I\left(\Delta \omega, \beta=\beta_{\mathrm{av}}\right)$. The latter three values are printed out normalized with respect to the asymptotic $\Delta \omega^{-5 / 2}$-wing.

As explained above the program STBRHY may be replaced by the program STBRHY listed at the end for the case of no lower state interactions. Both programs are written in identically the same manner with the only difference that the program STBRHY listed at the end has a simpler matrix ordering and a much simpler and faster calculation of the $K$-matrix according to eq (27), which makes the provision to read in the $K$-matrix no longer desirable.

PROGRAM STBRHY
C
MENSION DQQ(30), PFAC(6),STRONG(34),FF(1100),SJUU(136,31)
$1 \operatorname{SJLL}(6,5), \operatorname{RDM}(16,3), \operatorname{SLUU}(16), \operatorname{SLLL}(3), \operatorname{NQUU}(34,2), \operatorname{NQLL}(34,2)$,
$2 \operatorname{MMUU}(34,2), \operatorname{MMLL}(34,2), \operatorname{NQCU}(30), \operatorname{NQCL}(30), \operatorname{NXX}(30), \operatorname{NST}(30)$,
$3 \operatorname{DST}(30), W(300,5)$
COMMON/FDAT/P1, P2,B1, A2, B2, PPFF, CCREAL, CCIMAG
COMMON/PFW/FIELD (301)
COMMON/11/SLOP $(595,22,2)$
COMMON/22/DIPOL (34,2), NUR(2), IPI, NCQ, $\operatorname{BET}, \operatorname{FPAR}(6,30), \operatorname{ADIFF}(34,2)$,
1 NSTEP
COMMON/33/AMATR $(34,68), \operatorname{BMATR}(34,34), \operatorname{CMATR}(34,34)$
EQUIVALENCE (SJUU, AMATR), (FF (601), NQUU), (FF (669), NQLL),
1 (FF (737), MMUU), (FF (805), MMLL), (FF (873), NQCU), (FF (903), NQCL),
2 (FF (933), SJLL) , (FF (963) , RDM) , (FF (1011), SLUU) , (FF (1027), STRONG)
$\operatorname{FIELD}(1)=0.0$
READ 100, ( $(W(I, J), J=1,5), \quad I=1,300)$
100 FORMAT (5E12.4)
120 READ 150, DEN,TEMP,NNUU,NNLL,GIN,DGG,NTOT,NFAC, (PFAC (I) , I=1, 6)
150 FORMAT (2E10.2,I3,I2,2F10.2,2I5/6F10.5)
IF (EOF,60) 577, 170
C CALCULATION OF MICROFIELD DISTRIBUTION FUNCTION
170 SHIELD $=0.0898 * \operatorname{DEN} * *(1 . / 6$.$) / SQRT (TEMP)$
DO $161 \mathrm{~J}=1,5$
$161 \operatorname{DST}(\mathrm{~J})=0.2 * \operatorname{FLOAT}(\mathrm{~J}-1)$
DO $163 \mathrm{I}=1,300$
DO $162 \mathrm{~J}=1,5$
$162 \mathrm{DQQ}(\mathrm{J})=\mathrm{W}(\mathrm{I}, \mathrm{J})$
CALL POLY5 (DST,DQQ,5,STRONG,SLUU,SHIELD,0.1,1)
$163 \operatorname{FIELD}(\mathrm{I}+1)=\operatorname{SLUU}(1)$
C CALCULATION OF AVERAGE IONFIELD
DO 111 I $=1,300$
$\mathrm{AM}=\mathrm{I}$
$111 \mathrm{FF}(\mathrm{I})=\operatorname{FIELD}(\mathrm{I}+1)$ * $\mathrm{AM} / 10$.
CALL WEDDLE (0.1, 300, FF, BAV, 0.)
$D Y=1 . / 300 . \quad \$ \quad Y=0$.
DO $113 \mathrm{I}=1,10$
$\mathrm{Y}=\mathrm{Y}+\mathrm{DY} \quad \$ \mathrm{~B}=1 . / \mathrm{Y}$
$113 \mathrm{FF}(\mathrm{I})=\mathrm{WFLD}(\mathrm{B}) * \mathrm{~B}^{* *} 3$
CALL WEDDLE (DY, 10, FF, DB, 0.)
$B A V=B A V+D B$
READ 175,NPUNCH, (NUR (I), I=1,2),IPI,NNCCQ
175 FORMAT (5I10)
C FOR NPUNCH GT 1 K-MATRIX IS READ IN
C FOR NPUNCH LT 1 K -MATRIX IS CALCULATED
C FOR NPUNCH EQ 1 K-MATRIX IS CALCULATED AND PUNCHED OUT
IF (NPUNCH.LE.l) GO TO 178
DO 173 MBLOCK $=1,2$
NUPP $=$ NUR(MBLOCK)
NRRAB $=($ NUPP +1$) *$ NUPP $/ 2$
DO $174 \mathrm{~J}=\mathrm{IPI}, \mathrm{NNCCQ}$
READ 172, ( $\mathrm{FF}(\mathrm{I}), \mathrm{I}=1, \mathrm{NRRAB})$
172 FORMAT (5El6.9)
DO 171 I $=1, N R R A B$
$171 \operatorname{SLOP}(I, J, M B L O C K)=F F(I)$
174 CONTINUE
173 CONTINUE
C CALCULATION OF WAVELENGTH IN STANDARD AIR
178 AUU $=$ NNUU
ALL $=$ NNLL
SL1 $=109678.758 *(1 . /($ ALL*ALL $)-1 . /(A U U * A U U))$

```
    SL2 = SLl * SLl
        ALDD = 1.000064328+2949810./(1.46E+10-SL2)+25540./(4.1E+9-SL2)
        ALAM = 1.E+8/(ALDD * SLl)
        PRINT 180, DEN, TEMP, NNUU, NNLL, ALAM, SHIELD
    180
    FORMAT (lHl,* DENSITY =* ElO.2* TEMPERATURE =*ElO.2,
    l * QUANTUMNUMBERS N UPPER =*I3* N LOWER =*I2* WAVELENGTH =*
    2 F8.2* ANGSTR0M*/10X,*SHIELDING PARAMETER =*F6.3,10X,
    3 * INCLUDING LOWER STATE INTERACTION *//)
    IF (NNUU.LE.5.AND.NNLL.LE.3) GO TO 200
    IF (NNUU.LE.8.AND.NNLL.LE.2) GO TO }20
    IF (NNUU.LE.16.AND.NNLL.EQ.l) GO TO 200
    PRINT 190
    190 FORMAT (* OVERFLOW OF MATRICES, PROGRAM NOT EXECUTED*)
    577 CALL EXIT
C
C RADIAL MATRIXELEMENTS AND TOTAL LINESTRENGTH
    200 IPI = 2
        STOT = 0.
        DO 330 KLL = 1,NNLL
        LLL = KLL - 1 $ ALL = LLL
        DO 3l0 KUU = l,NNUU
        LUU = KUU - 1 $ AUU = LUU $ RRMM = 0.
        IF (IABS(KUU - KLL).NE.1) GO TO 310
        RRMM = RADMAT(NNUU,LUU,NNLL,LLL) * S3J(ALL,AUU,1.,0.,0.,0.)
        STOT = STOT + (2.*ALL+1.)*(2.*AUU+1.) * RRMM * RRMM
    310 RDM(KUU,KLL) = RRMM
    330 CONTINUE
    SQST = SQRT(STOT)
C
C TRANSFORMATION MATRICES
    DO 370 NKK = 1,2
    NN = NNLL
    IF (NKK.EQ.2) NN = NNUU
    ANN = NN $ ANl = 0.5 * (ANN - 1.)
    KQQ = 2 * NN - 1 $ NRUN = 0
    DO 370 KL = l,NN
    AL = KL - l
    FACC = SQRT(2.*AL+1.)
    IF(NKK.EQ.1) SLLL(KL) = FACC
    IF(NKK.EQ.2) SLUU(KL) = FACC
    DO 370 KM = 1,KL
    AM = KM - l $ NRUN = NRUN + l
    QQ = -ANN $ ALIM = ANN - 0.9 - AM
    DO 370 KQ = 1,KQQ
    QQ = QQ + 1. $ SIIJ = 0. $ ABQQ = ABS (QQ)
    IF (ABQQ.GT.ALIM) GO TO 360
    NQ = ABQQ + 0.l
    IF (MOD(NN+NQ+KM+1,2).NE.1) GO TO 360
    AMQ = 0.5* (AM - QQ) $ APQ = 0.5 * (AM + QQ)
    SIIJ = FACC * S3J(ANL,ANI,AL,AMQ,APQ,-AM)
    360 IF (NKK.EQ.l) SJLL(NRUN,KQ) = SIIJ
    IF (NKK.EQ.2) SJUU(NRUN,KQ) = SIIJ
    370 CONTINUE
C
C DIPOLMATRIXELEMENTS IN PARABOLIC STATES AND MATRIX ORDERING
    KQUU = 2 * NNUU - 1 $ KQLE = 2 * NNLL - 1
    DO 500 MBLOCK = 1,2
    AMLCK = MBLOCK $ MBL = MBLOCK - 1 $ AMBL = MBL
    NRUN = 0 $ NCQ = 0 $ NAQU = -NNUU
    DO 480 KUU = l,KQUU
    NAQU = NAQU + l $ KQU = NNUU + NAQU $ NAQL = -NNLL
    DO 480 KLL = l,KQLL
    NAQL = NAQL + l
```

```
    KQL = NNLL + NAQL
    NXDIF = NNUU * NAQU - NNLL * NAQL
    AXDIF = NXDIF
    IF (NXDIF.LT.0) GO TO 400
    NCQ = NCQ + l
    IF (MBLOCK.EQ.2) GO TO 400
    NQCU (NCQ) = NAQU $ NQCL(NCQ) = NAQL
    NXX(NCQ) = NXDIF $ DQQ(NCQ) = 0
400
    NPHAS = (NNUU + NNLL + MBL - NAQU - NAQL - 2)/2
    PHASE = (-1.)**MOD (NPHAS,2)
    KMINL = NNLL - IABS(NAQL) $ KMINU = NNUU - IABS(NAQU)
    KML = -KMINL $ MKML = 2 * KMINL - l
    DO 470 KAML = l,MKML
    KML = KML + I
    IF (MOD(NNLL+NAQL+KML,2).NE.l) GO TO 470
    KMU = KML + MBL
    IF (IABS(KMU).GE.KMINU) GO TO 470
    IF (MOD(NNUU+NAQU+KMU,2).NE.l) GO TO 470
    NRUN = NRUN + l
    NQUU(NRUN,MBLOCK) = NAQU $ NQLL(NRUN,MBLOCK) = NAQL
    MMUU (NRUN,MBLOCK) = KMU $ MMLL(NRUN,MBLOCK) = KML
    ADIFF(NRUN,MBLOCK) = AXDIF
    AMU = KMU $ AML = KML
    KBBU = IABS (KMU) + l $ KBBL = IABS (KML) + l
    RESULT = 0.
    DO 450 KKLU = KBBU,NNUU
    ALU = KKLU - l
    LLU = (KKLU-1) * KKLU / 2 + KBBU
    FACl = SLUU(KKLU) * SJUU(LLU,KQU)
    DO 450 KKLL = KBBL,NNLL
    IF (IABS(KKLU - KKLL).NE.I) GO TO 450
    ALL = KKLL - I
    LLL = (KKLL - l) * KKLL / 2 + KBBL
    FAC2 = SLLL(KKLL) * SJLL(LLL,,KQL)
    FACC = FAC2 * FACl * S3J (ALL,ALU,1.,AML,-AMU,AMBL)
    RESULT = RESULT + FACC * RDM(KKLU,KKLL)
4 5 0 ~ C O N T I N U E ~
    SIIJ = RESULT * PHASE / SQST
    DIPOL(NRUN,MBLOCK) = SIIJ
    IF (NXDIF.GE.O) DQQ(NCQ) = DQQ(NCQ) + AMLCK*RESULT*RESULT/STOT
    PRINT 425,NAQU,KMU,NAQL,KML,SIIJ
    425 FORMAT (16X,*(*,2I3,* I D I*,2I3,*) =*E16.8)
    4 7 0 ~ C O N T I N U E ~
    4 8 0 ~ C O N T I N U E ~
    NUR(MBLOCK) = NRUN
    PRINT }39
    393 FORMAT(/)
    500 CONTINUE
C
C REORDERING OF THE STARK COMPONENTS
    NCM1 = NCQ - l
    DO 333 II = l,NCM1
    IPP = II + l
    DO 333 JJ = IPP,NCQ
    IF (NXX(II).LE.NXX(JJ)) GO TO 333
    TEMPP = NXX(II) $ NXX(II) = NXX(JJ) $ NXX(JJ) = TEMPP
    TEMPP = DQQ(II) $ DQQ(II) = DQQ(JJ) $ DQQ(JJ) = TEMPP
    TEMPP = NQCU(II) $ NQCU(II) = NQCU(JJ) $ NQCU(JJ) = TEMPP
    TEMPP = NQCL(II) $ NQCL(II) = NQCL(JJ) $ NQCL(JJ) = TEMPP
333 CONTINUE
    PRINT 410, (NQCU(I),NQCL(I),NXX(I),DQQ(I), I = l,NCQ)
410 FORMAT (15X,* Q UPPER =*I3,* Q LOWER =*I3,* X =*I5,* DIPOL=*E16.8)
```

```
            PRINT 393
            JJ = 1 $ KMST = 0
    KMST = KMST + l
    DST(KMST) = 0.
    DO 325 II = JJ,NCQ
    IF (NXX(II).NE.NXX(JJ)) GO TO 325
    NST(KMST) = NXX(II) $ DST(KMST) = DST(KMST) + DQQ(II)
    J = II
325
    CONTINUE
    JJ = J + 1
    IF (JJ.GT.NCQ) GO TO 345
    G0 TO 3l5
    345 PRINT 355, (NST(I),DST(I),I = 1,KMST)
355 FORMAT (16X,*X =*I3,F20.10)
    PRINT }39
C
C K OPERATOR MATRIX
    IF (NPUNCH-1) 445,455,650
455 IPP = 2
    PUNCH 465,IPP,(NUR(I),I=1,2),IPI,NCQ,NNUU,NNLL
    FORMAT (5I10,15X,2I5)
    NTIME = KLOCK(0)
    DO 600 MBLOCK = 1,2
    MBL = 1 - MBLOCK $ AMBL = MBL $ NUPP = NUR(MBLOCK)
    DO 600 NRC = IPI,NCQ
    NCQU = NQCU(NRC) $ NCQL = NQCL(NRC)
    KCQU = NNUU + NCQU $ KCQL = NNLL + NCQL
    IMULIM = NNUU - IABS(NCQU) $ IMLLIM = NNLL - IABS(NCQL)
    DO 590 NRA = 1,NUPP
    NAQU = NQUU (NRA,MBLOCK) · $ KAMU = MMUU(NRA,MBLOCK)
    AMU = KAMU $ IABAMU = IABS (KAMU)
    NAQL = NQLL(NRA,MBLOCK) $ KAML = MMLL(NRA,MBLOCK)
    AML = KAML $ IABAML = IABS( KAML )
    KAQU = NNUU + NAQU $ KAQL = NNLL + NAQL
    KAAU = IABS (KAMU) + 1 $ KAAL = IABS (KAML) + 1
    DO 580 NRB = 1,NRA
    NRRAB = (NRA - 1) * NRA / 2 + NRB
    NBQU = NQUU (NRB,MBLOCK) $ KBMU = MMUU (NRB,MBLOCK)
    BMU = KBMU $ IABBMU = IABS (KBMU)
    NBQL = NQLL(NRB,MBLOCK) $ KBML = MMLL(NRB,MBLOCK)
    BML = KBML $ IABBML = IABS (KBML)
    KBQU = NNUU + NBQU $ KBQL = NNLL + NBQL
    KBBU = IABS (KBMU ) + 1 $ KBBL = IABS (KBML ) + 1
    NPHAS = MBL - NCQU - NCQL - (NAQU+NAQL+NBQU+NBQL)/2
    SIIJ = 0.
    DO 540 JLAL = KAAL,NNLL
    ALL = JLAL - l
    LALL = (JLAL - 1) * JLAL / 2 + IABAML + 1
    FACl = SJLL(LALL,KAQL)
    DO 540 JLAU = KAAU,NNUU
    ALU = JLAU - 1
    LALU = (JLAU - 1) * JLAU / 2 + IABAMU + 1
    FAC2 = SJUU(LALU,KAQU) $ FACl2 = FACl * FAC2
    DO 540 JLBL = KBBL,NNLL
    BLL = JLBL - 1
    LBLL = (JLBL - 1) * JLBL / 2 + IABBML + 1
    FAC3 = SJLL(LBLL,KBQL) $ FACl3 = FACl2 * FAC3
    MCLUP = MINO(JLBL,JLAL) - 1
    DO 540 JLBU = KBBU,NNUU
    BLU = JLBU - 1
    LBLU = (JLBU - l) * JLBU / 2 + IABBMU + 1
    FAC4 = SJUU(LBLU,KBQU) $ FACl4 = FACl3 * FAC4
    MCUUP = MINO(JLBU,JLAU ) - 1
```

```
    IUPP = MINO(IABS(JLAU+JLAL),IABS(JLBU+JLBL)) - I
    MCL = -IMLLIM - l
    DO 540 JMCLL = 1,IMLLIM
    MCL = MCL + 2 $ IABCML = IABS (MCL)
    IF (IABCML.GT.MCLUP) GO TO 540
    AMCL = MCL
    LALL = (JLAL - 1) * JLAL / 2 + IABCML + 1
    LBLL = (JLBL - l) * JLBL / 2 + IABCML + 1
    FAC5 = SJLL(LALL,KCQL) * SJLL(LBLL,KCQL)
    FACl5 = FACl4 * FAC5
    IF (ABS(FACl5).LT.l.E-8) GO TO 540
    MCU = -IMULIM - l
    DO 535 JMCUU = 1,IMULIM
    MCU = MCU + 2 $ IABCMU = IABS (MCU )
    IF (IABCMU.GT.MCUUP) GO TO 535
    AMCU = MCU
    LALU = (JLAU - 1) * JLAU / 2 + IABCMU + 1
    LBLU = (JLBU - 1) * JLBU / 2 + IABCMU + 1
    FAC6 = SJUU(LALU,KCQU) * SJUU(LBLU,KCQU)
    IF (ABS(FAC6).LT.l.E-8) G0 T0 535
    MGGL = MCL - MCU $ AMGL = MGGL
    ILOW = MAXO(MBL,IABS (MGGL),IABS (JLAU-JLAL),IABS(JLBU-JLBL)) + l
    IF (ILOW.GT.IUPP) GO TO 535
    FACl6 = FACl5 * FAC6 * ((-l.)**MOD(NPHAS-MGGL,2))
    DO 530 JLDD = ILOW,IUPP
    IF (MOD(JLAL+JLAU+JLDD,2).NE.0) GO TO 5l5
    IF (KAMU.EQ.O.AND.KAML.EQ.O) GO TO 530
    IF (MCL.EQ.O.AND.MCU.EQ.O) GO TO 530
515 IF (MOD(JLBL+JLBU+JLDD,2).NE.0) GO TO 520
    IF (KBMU.EQ.O.AND.KBML.EQ.O) GO TO 530
    IF (MCL.EQ.O.AND.MCU.EQ.O) GO TO 530
    FAC7 = S3J (ALL, ALU,ALDD , -AMCL, AMCU,AMGL) * (2.*ALDD + 1.)
    FAC7 = FAC7 * S3J (ALL,ALU,ALDD,-AML,AMU,AMBL)
    FAC7 = FAC7 * S3J (BLL,BLU,ALDD,-AMCL,AMCU,AMGL)
    FACC = FAC7 * S3J(BLL,BLU,ALDD,-BML,BMU,AMBL) * FACl6
    SIIJ = SIIJ + FACC
530
5 3 5 ~ C O N T I N U E ~
540 CONTINUE
    SLOP(NRRAB,NRC,MBLOCK) = SIIJ
580 FF(NRRAB) = SIIJ
5 9 0 ~ C O N T I N U E ~
    IF (NPUNCH.EQ.l) PUNCH 630,(FF(I),I=l,NRRAB)
630 FORMAT (5El6.9)
6 0 0 ~ C O N T I N U E ~
    NTIME = -NTIME + KLOCK(0)
    PRINT 617, NTIME
    617 FORMAT (16X,*COMPUTERTIME FOR CALCULATING K-OPERATOR MATRIX =*I8/)
C
C BASIC CONSTANTS AND ARRAY FOR G-FUNCTION CONSTANTS
    650
    SDEN = SQRT(DEN)
    FAC = 2064.936 * TEMP * SQRT(TEMP/DEN)
    CFAC = 4.5645E-7 * SDEN/TEMP
    DEBROG = 2.1027E-6/SQRT(TEMP)
    ANN = NNUU
    RMIN = DEBROG + ANN * ANN * 7.9376E-9
    BET = 5.6558E-5 * DEN**(1./6.)
    ASY = 0.0
    DO 270 NRC = IPI,NCQ
    QC = NXX(NRC) $ C = CFAC * QC
    Pl = -l.671086 * FAC * C * SQRT(C)
    BS = 3. * QC * DEBROG / RMIN
```

```
    STRONG(NRC) = 0.269-2.*(((1.-COSF(BS))/BS+SINF(BS ))/BS-COSINT(BS ))
    PPFF = -l.128379 * FAC * C * C
    P2 = PPFF * (STRONG(NRC)-2.*LOGF (2.*C))
    FPAR(1,NRC) = P1 $ FPAR(2,NRC) = P2
    FPAR(3,NRC) = 0.5 * (P2/Pl)**2 $ FIN = LOGF(QC)
    A2 = P2*((PFAC (3)*FIN+PFAC (2))*FIN+PFAC(1))
    B2 = (PFAC (6)*FIN+PFAC(5))*FIN + PFAC(4)
    IF (B2.LT.0.) A2 = 0.
    FPAR(4,NRC) = A2 $ FPAR(5,NRC) = B2 $ FPAR(6,NRC) = PPFF
220 FORMAT (/13X*Pl*18X*P2*18X*B1*18X*A2*18X*B2*17X*STRONG*/)
    PRINT 240,((FPAR(K,I),K = l,5),STRONG(I), I = IPI,NCQ)
240 FORMAT (6E20.4)
    PRINT 280, FAC, CFAC, BET, ASY, DEBROG, BAV, NFAC
280 FORMAT(/* FAC =*Ell.4,* CFAC =*Ell.4,* BET =*Ell.4,* ASY =*Ell.4,
        * DEBROG =*Ell.4,* BAV =*F7.4,* INTEGRATIONFACTOR =*I2//
    2 5X*DOM*8X*DLAM*8X*ITOT*6X*IHOLTS*8X*ASY*1OX*WING*7X*WHOLTS*7X,
    3 *WWOO*8X*WWBB*8X*WWNG*8X*TIME STEPS*/)
    ADLFAC = 4.23538E-15 * SDEN * ALAM * ALAM
C
C CALCULATION OF THE IONFIELD INTEGRAL
    N12 = 12 * NFAC $ AN12 = N12 $ N30 = 30 * NFAC $ AN30 = N30
    DMCRT = BET * BAV * NXX(NCQ) * 5.
    G = GIN - DGG
    DO 950 MM = l,NTOT
    NTIME = KLOCK(0)
    NSTEP = O $ G = G + DGG
    DOM = 10. ** G $ DLAM = ADLFAC * DOM
    WING52 = -0.2992067103 * ASY/(SQRT(DOM) * DOM * DOM)
    FHOLTS = 0. $ AWING = 0.
    DO 815 NRC = IPI,NCQ
    Pl = FPAR(1,NRC) $ P2 = FPAR(2,NRC) $ Bl = FPAR(3,NRC)
    A2 = FPAR (4,NRC) $ B2 = FPAR (5,NRC) $ PPFF = FPAR(6,NRC)
    CALL FOUTR(DOM)
    AWING = AWING + DQQ(NRC) * CCIMAG * 2./(DOM*DOM)
    QC = NXX(NRC) $ BETFAC = BET * QC $ BCRIT = DOM/BETFAC
815 FHOLTS = FHOLTS + DQQ(NRC) * WFLD(BCRIT) / BETFAC
    IF (DOM.GT.DMCRT) GO TO 985
    AIRES = 0.
    IF(DOM.GT.(-3.*P2)) GO TO 840
    ANQ1 = NXX(NCQ)
    BCRIT = (DOM - P2)/(ANQ1*BET)
    DB = BCRIT/AN12 $ B = 0.
    DO 820 J = l,Nl2
    B = B + DB
820 FF(J) = AIIM(DOM,B) * WFLD(B)
    CALL WEDDLE (DB,N12,FF,AIII,0.)
    AIRES = AIII
    DY = 1./(BCRIT*AN30) $ Y = 0.
    DO 830 J = l,N30
    Y = Y + DY $ B = 1./Y
830 FF(J) = B * B * AIIM(DOM,B) * WFLD(B)
    CALL WEDDLE (DY,N30,FF,AIII,0.)
    AIRES = AIRES + AIII
    GO TO 980
840 BCRCR = DOM/BET
    EPSPS = -P2/BET
    DO 957 NQ = IPI,KMST
    ANQ = NST(NQ)
    BCR = BCRCR/ANQ $ EPS = EPSPS/ANQ
    IF (NQ.EQ.IPI) GO TO 907
    SLl = 1./(GAM - BCR)
```

G0 T0 908
$907 \mathrm{SLl}=0$.
908 SL2 $=1 . / E P S$
$\mathrm{SL} 3=1 . /(\mathrm{BCR}+\mathrm{EPS}) \quad \$ \mathrm{SL} 4=1 . /(\mathrm{BCR}-\mathrm{EPS})$
IF (NQ.EQ.KMST) GO TO 911
ANQ1 $=\mathrm{NST}(\mathrm{NQ}+1)$
GAM $=0.5 *($ BCR-EPS $+($ BCRCR + EPSPS $) / \mathrm{ANQ} 1)$
G0 TO 912
$911 \mathrm{GAM}=0.5 *(\mathrm{BCR}-\mathrm{EPS})$
912 SL5 $=1 . /(\mathrm{BCR}-\mathrm{GAM})$
CRIT = SL2 - SL5
$\mathrm{Y}=\mathrm{SLl}$
IF (NQ.EQ.IPI) GO TO 913
$B=B C R+1 . / Y$
$F A=\operatorname{AIIM}(D O M, B) * W F L D(B) /(Y * Y)$
G0 TO 914
$913 \mathrm{FA}=0$.
914 DY $=($ SL2 - SLl $) /$ AN12
DO $917 \mathrm{~J}=1, \mathrm{~N} 12$
$Y=Y+D Y \quad \$ \quad Y l=1 . / Y \quad \$ \quad B=B C R+Y l$
$917 \mathrm{FF}(\mathrm{J})=\mathrm{Yl}$ * Yl * AIIM(DOM,B) * WFLD(B)
CALL WEDDLE (DY,N12,FF,AIII,FA)
AIRES $=$ AIRES + AIII
$Y=\operatorname{SL} 3 \quad \$ \quad B=1 . / Y$
$F A=B * B * A I I M(D O M, B) * W F L D(B)$
$D Y=($ SL4 - SL3 $) /$ AN12
DO $927 \mathrm{~J}=\mathrm{l}, \mathrm{Nl2}$
$Y=Y+D Y \quad \$ \quad B=1 . / Y$
$927 \mathrm{FF}(\mathrm{J})=\mathrm{B} * \mathrm{~B}$ * $\operatorname{AIIM}(\mathrm{DOM}, \mathrm{B}) * \operatorname{WFLD}(\mathrm{~B})$
CALL WEDDLE (DY,N12,FF,AIII,FA)
AIRES = AIRES + AIII
IF (CRIT.LE.O.) GO TO 977
$Y=\operatorname{SL5} \$ B=B C R-1 . / Y$
$F A=\operatorname{AIIM}(D O M, B) \quad * \operatorname{WFLD}(B) /(Y * Y)$
DY $=$ CRIT/AN12
DO $937 \mathrm{~J}=\mathrm{l}, \mathrm{Nl2}$
$Y=Y+D Y \quad \$ \quad Y l=1 . / Y \quad \$ \quad B=B C R-Y l$
$937 \mathrm{FF}(\mathrm{J})=\mathrm{Yl}$ * Y1 * AIIM(DOM,B) * $\operatorname{WFLD}(\mathrm{B})$
CALL WEDDLE (DY,N12,FF,AIII,FA)
AIRES = AIRES + AIII
957 CONTINUE
IF (GAM.LT. 5.) GO TO 968
$\mathrm{Y}=1 . / \mathrm{GAM} \$ \quad \mathrm{DY}=(0.2-\mathrm{Y}) /$ AN12
FA $=$ GAM * GAM * AIIM(DOM, GAM) * WFLD (GAM)
DO $967 \mathrm{~J}=\mathrm{l}, \mathrm{N} 12$
$Y=Y+D Y \quad \$ \quad B=1 . / Y$
$967 \mathrm{FF}(\mathrm{J})=\mathrm{B} * \mathrm{~B} * \mathrm{AIIM}(\mathrm{DOM}, \mathrm{B}) * \operatorname{WFLD}(\mathrm{~B})$
CALL WEDDLE (DY, N12,FF,AIII,FA)
AIRES = AIRES + AIII
SL4 $=0.2$
GO TO 977
968 SL4 $=1 . / \mathrm{GAM}$
$977 \mathrm{~B}=0$.
$\mathrm{DB}=1 . /\left(\mathrm{SL} 4{ }^{*}\right.$ AN30)
DO $947 \mathrm{~J}=1 . \mathrm{N} 30$
$B=B+D B$
$947 \mathrm{FF}(\mathrm{J})=\mathrm{AIIM}(\mathrm{DOM}, \mathrm{B}) * \operatorname{WFLD}(\mathrm{~B})$
CALL WEDDLE (DB, N30,FF,AIII, O.)
AIRES $=$ AIRES + AIII
980 WWBB $=($ AIIM $($ DOM, BAV $)+$ FHOLTS $) /$ WING52
G0 TO 990
985 AIRES $=$ AIIM $($ DOM, BAV $)+$ FHOLTS
$W W B B=0$.
WING = AIRES/WING52
WINHOL = FHOLTS/WING52
WWOO = (AIIM(DOM, 0.) + FHOLTS)/WING52
WWNG = (AWING + FHOLTS)/WING52
NTIME = -NTIME + KLOCK(0)
950 PRINT 978,DOM,DLAM,AIRES,FHOLTS,WING52,WING,WINHOL,WWOO,WWBB,
l WWNG,NTIME,NSTEP
978 FORMAT (1OE12.4,IlO,I6)
GO TO 120
END
C
C
FUNCTION AIIM(DOM,B)
C
C CALCULATION OF I(DOM,B) FOR THE GENERAL CASE INCLUDING LOWER
C
220 DIIF(NRC) = CCIMAG
DO 700 NRA = l,NUPP
IF (NRA.LE.NRB) NRRAB = (NRB - l) * NRB / 2 + NRA
IF (NRA.GT.NRB) NRRAB = (NRA - l) * NRA / 2 + NRB
ARRR = 0. \$ AIII = 0.
DO 600 NRC = IPI,NCQ
TEMP = SLOP(NRRAB,NRC,MBL)
ARRR = ARRR + TEMP * DRRF (NRC)
600 AIII = AIII + TEMP * DIIF (NRC)
AIII = AIII * 6.2831853072
AMATR(NRB,NRA) = AIII \$ CMATR(NRB,NRA) = AIII
7 0 0 ~ B M A T R ( N R B , N R A ) ~ = ~ A R R R ~ * ~ 6 . 2 8 3 1 8 5 3 0 7 2 ~
750 BMATR(NRB,NRB) = BMATR(NRB,NRB) + DOMRB
CALL ENVERS (NUPP)
DO 320 NRB = 1,NUPP
DO 320 NRA = 1,NUPP
TEMP = 0.
DO 300 NRC = 1,NUPP
TEMP = TEMP + BMATR(NRB,NRC) * AMATR(NRC,NRA+NUPP)
320 FMATR (NRB,NRA) = TEMP
DO 420 NRB = 1,NUPP
DO 420 NRA = 1,NUPP
TEMP = 0.
DO 400 NRC = l,NUPP
400 TEMP = TEMP + FMATR(NRB,NRC) * BMATR (NRC,NRA)
420 AMATR(NRB,NRA) = TEMP + CMATR(NRB,NRA)
CALL ENVERS (NUPP)
DO 795 NRB = l,NUPP
DO }795\mathrm{ NRA = 1,NUPP
795 AIIM = AIIM+AML*DIPOL(NRB,MBL)*DIPOL(NRA,MBL)**AMATR(NRB,NRA+NUPP)
800 CONTINUE

```

AIIM \(=\) AIIM * 0.3183099 RETURN
END
C
C
C
SUBROUTINE POLY5 (X,Y,NUMX,XN,YN,XNI,DX,NXNUM)
A 5 POINT POLYNOMIAL INTERPOLATION ROUTINE
FOR EACH POINT TO BE INTERPOLATED8 THE 5 NEAREST KNOWN POINTS
ARE CHOSEN, AND A 5TH DEGREE POLYNOMIAL IS FITTED TO THESE POINTS X AND Y ARE THE ARRAYS OF KNOWN POINTS ON THE CURVE
NUMX IS THE NUMBER OF KNOWN POINTS
two new arrays, Xn and yn will be generated for the nxnum values
OF X AND \(\mathrm{F}(\mathrm{X})\) STARTING WITH X=XNI IN INCREMENTS OF DX
DIMENSION X(5),Y(5),XN(5),YN(5)
\(\mathrm{J}=2\)
DO 8 I=l, NXNUM
EYE=I-1
XN(I) \(=\mathrm{XNI}+\mathrm{DX} * \mathrm{E}_{\mathrm{E}}\)
IF (XN(I)-X(1)) 1,2,3
\(\mathrm{YN}(\mathrm{I})=\mathrm{Y}(\mathrm{l})\)
GO T0 8
3 IF (XN(I).LE.X(J)) GO TO 4
IF (J.GE.NUMX) GO TO 7
\(\mathrm{J}=\mathrm{J}+\mathrm{l}\)
G0 T0 3
\(1 \quad \mathrm{~L}=1\)
GO T0 70
7 L = NUMX - 4
GO TO 70
\(\mathrm{L}=\mathrm{J}-\mathrm{l}\)
IF (J.GT.2) L = J - 2
IF (J.GT.3) L = J - 3
IF ((J+l). GT. NUMX) L = J - 4
\(\mathrm{YN}(\mathrm{I})=0.0\)
LLL=L+4
DO \(75 \mathrm{~K}=\mathrm{L}, \mathrm{LLL}\)
TERM=1. 0
DO \(74 \mathrm{M}=\mathrm{L}, \mathrm{LLL}\)
IF (K.EQ.M) GO TO \(\mathrm{h}_{4}\)
TDEN \(=\mathrm{X}(\mathrm{K})-\mathrm{X}(\mathrm{M})\)
TNUM \(=\mathrm{XN}(\mathrm{I})-\mathrm{X}(\mathrm{M})\)
TERM=TERM*TNUM/TDEN
74 CONTINUE
TERM \(=\mathrm{Y}(\mathrm{K}) *\) TERM
\(75 \mathrm{YN}(\mathrm{I})=\mathrm{YN}(\mathrm{I})+\) TERM
8 CONTINUE
RETURN
END
C
C
SUBROUTINE FOUTR (DOM)
C
C
FOURIERTRANSFORM OF THERMAL AVERAGE FOR UNIFIED THEORY
COMMON/FDAT/P1, P2,B1, A2, B2, PPFF, CCREAL, CCIMAG
ARG \(=\operatorname{ABSF}(\mathrm{DOM})\)
\(\mathrm{Z}=\mathrm{Bl}{ }^{*} \mathrm{ARG}\)
IF (Z.LE.O.OO1) GO TO 600
AR2 \(=\) ARG * ARG
IF (Z.LE.40.) GO TO 300
\(\mathrm{FACl}=-0.2992067103\) * Pl/(SQRT(ARG) * AR2)
\(\mathrm{CC}=\mathrm{FACl} *((1 .-1.3125 / \mathrm{Z}) * 0.625 / \mathrm{Z}+1\).
SS = CC - FACl * \(1.25 / \mathrm{Z}\)
GO TO 500
CALL BSJYOl (Z, AJO, YO, AJl, Yl)
```

    FACl = Yl/(2.*Z) + AJl - Y0
    FAC2 = AJO + Yl - AJl/(2.*Z)
    CINE = COSF(Z)
    SINE = SINF(Z)
    FCC = P2 * Bl * Bl
    CC = FCC * (CINE*FACl+SINE*FAC2)
    SS = FCC * (CINE*FAC2-SINE*FACl)
    IF (A2.EQ.O.) GO TO 500
    Z = B2 * ARG
    IF (Z.GT.l0.) GO TO 400
    CALL BSJYOl (Z, AJO, YO, AJl, Yl)
    FACl = ((AJl-Y0)*16.*Z-36.*AJO-28.*Y1)*Z+15.*Y0-3.*AJl
    FAC2 = ((AJO+Y1)*16.*Z-36.*YO+28.*AJl)*Z-15.*AJO-3.*Y1
    CINE = COSF(Z)
    SINE = SINF(Z)
    FCC = A2 * B2/6.
    CC = CC + FCC * (CINE*FACl+SINE*FAC2)
    SS = SS + FCC * (CINE*FAC2-SINE*FACl)
    GO TO 500
    400 FACl = 0.1322319336 * A2 * B2 * Z** (-3.5)
    FCC = FACl * (1. - (3.9375/Z + 1.)*4.375/Z)
    CC = CC + FCC
    SS = SS - FCC - FACl*8.75/Z
    500 IF (DOM.LT.O.) SS = -SS
    CCREAL = -AR2 * SS
    CCIMAG = AR2 * CC
    RETURN
    6 0 0
    CCREAL = 0.3183099 * (P2*Bl-A2) * DOM
    CCIMAG = -0.3183099 * P2
    RETURN
    END
    C
C
SUBROUTINE ENVERS (NN)
C
COMMON/33/C (34,68), BMATR (34,34), CMATR (34, 34)
N = NN
IF (N.GT.l) GO TO 10
C(1,2) = 1./C(1,1)
RETURN
1 0
DO 300 K = l,N
KN = K + N
NP = KN - I
KPl = K + l
TEMP = 1./C (K,K)
DO 100 I = KPl,NP
100 C(K,I) = C(K,I) * TEMP
C(K,KN) = TEMP
DO 300 L = l,N
IF (K.EQ.L) GO TO 300
TEMP = C(L,K)
DO 200 I = KPl,NP
200 C(L,I) = C(L,I) - C(K,I) * TEMP
C(L,KN) = -C(K,KN) * TEMP
300 CONTINUE
RETURN
END
C
C
FUNCTION RADMAT(N1,L1,N2,L2)
C
DIMENSION F(2)
RADMAT = 0.0

```
```

    IF(IABS(Ll - L2).NE.l) RETURN
    NA = N1 $ NB = N2 $ L = Ll
    IF(Ll.GT.L2) GO TO 10
    NA = N2 $ NB = N1 $ L = L2
    10 ANA = NA
    BNB = NB $ AL = L
    IF(N1.EQ.N2) GO TO 60
    RADMAT = FCTRL(ANA + AL) * FCTRL(BNB + AL - 1.)
    RADMAT = RADMAT/(FCTRL(ANA - AL - l.) * FCTRL(BNB - AL))
    RADMAT = SQRT(RADMAT)
    I = BNB + AL $ AL = 2. * AL
    RADMAT = RADMAT * ((-1.0)**XMODF(I,2))/(4.*FCTRL(AL - l.))
    ABAB = (ANA - BNB)/(ANA + BNB)
    AB2 = 1./(ANA - BNB)**2
    RADMAT = RADMAT * ABAB**(NA+NB)*(4.*ANA*BNB*AB2)**(L+l)
    Z = -4. * ANA * BNB * AB2
    NR = NA - L - l $ NRP = NB - L $ ANRP = NRP
    DO 30 MM = 1,2
    F(MM) = 1.0
    IF(MM.EQ.2) NR = NR + 2
    K = MINO (NR,NRP)
    IF(K.EQ.O) GO TO 30
    ANR = NR $ PROD = 1.0
    DO 40 M = 1,K
    AM = M - l
    PROD = PROD * (AM-ANR)*(AM-ANRP)*Z/((AL+AM)* (AM+1.))
    40 F(MM) = F(MM) + PROD
    30 CONTINUE
    RADMAT = RADMAT * (F(1) - ABAB * ABAB * F(2))
    RETURN
    60 RADMAT = -1.5 * ANA * SQRT((ANA+AL)*(ANA-AL))
RETURN
END
c
C
FUNCTION FCTRL(A)
c
C CALCULATION OF FACTORIALS
DIMENSION FCTI(20)
DATA ((FCTI(I),I=l,20) =1.0,2.0,6.0,24.0,120.0,720.0,5040.0,
l 40320.0,362880.0,3628800.0,39916800.0.479001600.0,
2 6227020800.0,87178291200.0,1307674368000.0,
3 2.0922789888E13, 3.55687428096E14, 6.402373705728E15,
4 l.2164510040883 El7, 2.4329020081766 El8)
I = A + 0.I
IF(I) 50,60,70
50 FCTRL = 0.0
RETURN
60 FCTRL = 1.0
RETURN
70 IF (I.GT.20) GO TO l30
FCTRL=FCTI(I)
RETURN
130 F=20.0
FCTRL=FCTI (20)
D0 l3l J=2l,I
F=F+l.0
131 FCTRL=FCTRL*F
RETURN
END
C
C
FUNCTION S3J (FJl, FJ2, FJ3, FM1, FM2, FM3)

```

CALCULATION OF 3J-SYMB0L, FJl+FJ2+FJ3 HAS TO BE LESS THAN 38 DIMENSION GAM(39)
DATA ( \((\mathrm{GAM}(\mathrm{I}), \mathrm{I}=1,39)=1 ., 1 ., 2 ., 6 ., 24 ., 120 ., 720 ., 5040 ., 40320 .\),
1 362880.,3628800.,39916800.,479001600.,6227020800.,87178291200.,
2 1307674368000.,2.0922789888E+13,3.5568742810E+14,
\(36.4023737057 \mathrm{E}+15,1.2164510041 \mathrm{E}+17,2.4329020082 \mathrm{E}+18\),
\(4 \quad 5.1090942172 \mathrm{E}+19,1.1240007278 \mathrm{E}+21,2.5852016739 \mathrm{E}+22\),
\(5 \quad 6.2044840173 \mathrm{E}+23,1.5511210043 \mathrm{E}+25,4.0329146113 \mathrm{E}+26\),
\(6 \quad 1.0888869450 \mathrm{E}+28,3.0488834461 \mathrm{E}+29,8.8417619937 \mathrm{E}+30\),
\(7 \quad 2.6525285981 \mathrm{E}+32,8.2228386542 \mathrm{E}+33,2.6313083693 \mathrm{E}+35\),
\(88.6833176188 \mathrm{E}+36,2.9523279904 \mathrm{E}+38,1.0333147966 \mathrm{E}+40\),
\(93.7199332679 \mathrm{E}+41,1.3763753091 \mathrm{E}+43,5.2302261747 \mathrm{E}+44\) )
S3J=0.0
\(\mathrm{FJA}=\mathrm{FJl}+0.001\) \$ \(\mathrm{FJB}=\mathrm{FJ} 2+0.001 \$ \mathrm{FJC}=\mathrm{FJ} 3+0.001\)
IF (FJA.LT.0.0.OR.FJB.LT.0.0.OR.FJC.LT.O.O) RETURN
IF (AMOD (FJA,0.5).GT.0.002) RETURN
IF (AMOD (FJB,0.5).GT.0.002) RETURN
IF (AMOD (FJC, 0.5).GT.0.002) RETURN
\(F M A=F M 1 \$ F M B=F M 2 \$ F M C=F M 3\)
IF (ABS (FMA).GT.FJA.OR.ABS (FMB).GT.FJB.OR.ABS (FMC).GT.FJC) RETURN
\(B=F J A-F M A \$ P=F J B-F M B \$ D=F J C-F M C \$ G=F J A+F J B-F J C\)
IF ( \((\mathrm{ABS}(\mathrm{FMA}+\mathrm{FMB}+\mathrm{FMC})+\mathrm{AMOD}(\mathrm{B}, 1)+.\operatorname{AMOD}(\mathrm{P}, 1)+.\mathrm{AMOD}(\mathrm{D}, 1)+.\mathrm{DIM}(0 ., \mathrm{G})\)
1 +DIM(ABS (FJA-FJB),FJC)).GT.0.01) RETURN
\(I H H=F J A+F J B+F J C+2\).
\(I A A=F J B+F J C+F M A+1 . \$ I C C=-F J A+F J B+F J C+1\).
\(I F F=F J A-F J B+F J C+1 . \$\) IGG \(=G+1\).
\(J 1 M=B+1 . \$ J 2 M=P+1 . \$ J 3 M=D+1\).
\(J 1 P=F J A+F M A+1 . \$ J 2 P=F J B+F M B+1 . \$ J 3 P=F J C+F M C+1\).
\(E 2=\operatorname{GAM}(J 1 M) * G A M(J 1 P) * G A M(J 2 M) * G A M(J 2 P)\)
El \(=(\) GAM \((I C C) * \operatorname{GAM}(I F F) / G A M(I H H)) * G A M(I G G) * G A M(J 3 M) * G A M(J 3 P)\)
El \(=\) SQRT(E2/El)
\(E=F J A-F J B+F M C\)
IAEE \(=\mathrm{ABS}(\mathrm{E})+0.001\)
IEE \(=\) IAEE \(+1 \$\) Il \(=0\)
IF (E.GE.O.) GO TO 100
\(\mathrm{IEE}=-\mathrm{IAEE}+1\) \$ II = IAEE
100 I2 = ICC - 1
IF (J3M.LT.ICC) I2 = J3M -1
DO 150 I=I1,I2
\(\mathrm{E} 2=\mathrm{GAM}(\mathrm{I}+1) * \operatorname{GAM}(\mathrm{ICC}-\mathrm{I}) * \operatorname{GAM}(J 3 \mathrm{M}-\mathrm{I}) / \mathrm{GAM}(\) IAA -I\()\)
150 S3J \(=\) S3J \(+\left(\left((-1 .){ }^{* *}\left(\mathrm{I}-(\mathrm{I} / 2)^{*} 2\right)\right) / \mathrm{E} 2\right)^{*} \mathrm{GAM}(\mathrm{J} 1 \mathrm{M}+\mathrm{I}) / \mathrm{GAM}(\mathrm{IEE}+\mathrm{I})\)
\(I=A B S(F J A+F M B-F M C)+0.01\)
S3J \(=\) S3J \({ }^{*}\left((-1.){ }^{* *}(I-(I / 2) * 2)\right) / E 1\)
RETURN
END
C
C
SUBROUTINE WEDDLE (DX, N, F, A, FO)
C
C INTEGRATION SUBROUTINE
DIMENSION F(N)
\(A=0.0 \quad \$ \quad K=N-1\)
DO \(15 \mathrm{I}=1,6\)
SUM \(=0.0\)
D0 \(6 \mathrm{~J}=\mathrm{I}, \mathrm{K}, 6\)
6 SUM \(=\) SUM \(+F(J)\)
G0 T0 (8, 10, 12, 10, 8, 14), I
\(8 \mathrm{~A}=\mathrm{A}+5.0\) * SUM
GO TO 15
\(10 \mathrm{~A}=\mathrm{A}+\mathrm{SUM}\)
G0 TO 15
\(12 \mathrm{~A}=\mathrm{A}+6.0\) * SUM
```

            GO TO 15
    14 A = A + 2.0 * SUM
    15 CONTINUE
        A = 0.3 * DX * (A + FO + F(N))
        RETURN
    END
    C
C
FUNCTION WFLD(B)
C
C CALCULATION OF THE ION MICROFIELD DISTRIBUTION FUNCTION USING A
C 5POINT INTERPOLATION FOR THE DATA READ INTO THE MAINPROGRAM
COMMON/PFW/FIELD(301)
WFLD = 0.0
IF (B.LE.30.0) GO TO 200
SBS = 1./(B * SQRT (B))
WFLD = ((21.6 * SBS + 7.639) * SBS + 1.496) * SBS/B
RETURN
200 IF (B.LE.O.O) RETURN
J = (B + 0.2) * 10.0
L = J - I
IF (J.GT.2) L = J - 2
IF (J.GT.3) L = J - 3
IF (J.GT.300) L = 297
70 LLL = L + 4
DO 75 K = L,LLL
AK = K - l
TERM = 1.0
DO 74 M = L,LLL
IF (K.EQ.M) GO TO 74
AM = M - l
TERM = TERM * (10.*B - AM)/(AK - AM)
74 CONTINUE
TERM = TERM * FIELD (K)
75 WFLD = WFLD + TERM
RETURN
END
C
C
FUNCTION COSINT(X)
C
C CALCULATION OF THE COSINE INTEGRAL
TYPE DOUBLE Y2,PROD,SUM,PT,DK
IF(X.LE.O.) GO TO 50
X2 = X * X
IF(X.GT.20.) G0 T0 30
Y2 = DBLE(X2) \$ PROD = -Y2 * 0.5 \$ SUM = PROD * 0.5
DO 10 K = 2,50
DK = 2 * K
PROD = -PROD * Y2/(DK*(DK - 1.))
SUM = SUM + PROD/DK
PT = ABS(PROD * l.D+l0)
IF (ABS (SUM).GT.PT) G0 T0 20
10 CONTINUE
20 SS = SNGL(SUM)
COSINT = SS + 0.5772156649 + LOGF(X)
RETURN
30 FA = 1. \$ FB = 1. \$ PO = 1.
X2 = 1./X2
DO 40 K = 1,10
AK = 2 * K \$ PO = -PO * AK * X2 \$ FA = FA + PO
PO = PO * (AK + l.) \$ FB = FB + PO \$ PA = ABS (P0 * l.E+l0)
IF(PA.LE.FB) GO TO 45

```
\(45 \mathrm{FX}=\mathrm{FA} / \mathrm{X}\)
\(\mathrm{GX}=\mathrm{FB}\) * X 2
\(\operatorname{COSINT}=\mathrm{FX} * \operatorname{SIN}(\mathrm{X})-\mathrm{GX} * \operatorname{COS}(\mathrm{X})\)
RETURN
50 WRITE \((61,100)\) X
100 FORMAT (* X LESS OR EQUAL TO ZERO, \(X=\) *E17.9) RETURN
END

SUBROUTINE BSJYO1 (X, AJO, YO, AJI, Yl)
CALCULATION OF THE BESSEL FUNCTIONS JO, YO, JI, AND YI FOR AN ARGUMENT X
DIMENSION \(\mathrm{A}(7), \mathrm{B}(7), \mathrm{C}(7), \mathrm{D}(7), \mathrm{E}(7), \mathrm{F}(7), \mathrm{G}(7), \mathrm{H}(7)\)
DATA \(((A(I), I=1,7)=0.00021,-0.0039444,0.0444479\),
\(1 \quad-0.3163866,1.2656208,-2.2499997,1.0)\)
DATA ( \((\mathrm{B}(\mathrm{I}), \mathrm{I}=1,7)=-0.00024846,0.00427916,-0.04261214\),
\(1 \quad 0.25300117,-0.74350384,0.60559366,0.36746691)\)
DATA ( \((C(I), I=1,7)=0.00014476,-0.00072805,0.00137237\),
\(1-0.00009512,-0.0055274,-0.00000077,0.79788456)\)
DATA ( \((\mathrm{D}(\mathrm{I}), \mathrm{I}=1,7)=0.00013558,-0.00029333,-0.00054125\),
\(1 \quad 0.00262573,-0.00003954,-0.04166397,-0.78539816)\)
DATA ( \((\mathrm{E}(\mathrm{I}), \mathrm{I}=1,7)=0.00001109,-0.00031761,0.00443319\),
\(1 \quad-0.03954289,0.21093573,-0.56249985,0.5)\)
DATA \(((F) I), I=1,7)=0.0027873,-0.0400976,0.3123951\),
\(1-1.3164827,2.1682709,0.2212091,-0.6366198)\)
DATA ( \(\mathrm{G}(\mathrm{I}), \mathrm{I}=1,7)=-0.00020033,0.00113653,-0.00249511\),
\(1 \quad 0.00017105,0.01659667,0.00000156,0.79788456\) )
DATA \(((H(I), I=1,7)=-0.00029166,0.00079824,0.00074348\),
\(1 \quad-0.00637879,0.0000565,0.12499612,-2.35619449\) )
\(\mathrm{AX}=\operatorname{ABSF}(\mathrm{X})\)
IF (AX.GT.O.0) GO TO 10
\(\mathrm{AJO}=1 . \quad \$ \mathrm{YO}=-1 . \mathrm{E}+30 \quad \$ \quad \mathrm{AJl}=0 . \quad \$ \quad \mathrm{Yl}=-1 . \mathrm{E}+30\)
RETURN
10 IF (AX.GT.3.0) GO T0 50
\(X X=(A X / 3.0){ }^{* *} 2\)
\(\mathrm{AJO}=\mathrm{A}(\mathrm{l}) \quad \$ \quad \mathrm{YO}=\mathrm{B}(\mathrm{l}) \quad \$ \quad \mathrm{AJl}=\mathrm{E}(\mathrm{l}) \quad \$ \quad \mathrm{Yl}=\mathrm{F}(\mathrm{l})\)
DO \(20 \mathrm{M}=2,7\)
\(\mathrm{AJO}=\mathrm{AJO} * \mathrm{XX}+\mathrm{A}(\mathrm{M})\)
\(\mathrm{YO}=\mathrm{YO}{ }^{*} \mathrm{XX}+\mathrm{B}(\mathrm{M})\)
\(\mathrm{AJl}=\mathrm{AJI} * \mathrm{XX}+\mathrm{E}(\mathrm{M})\)
\(\mathrm{Yl}=\mathrm{Yl}{ }^{*} \mathrm{XX}+\mathrm{F}(\mathrm{M})\)
AJl \(=\) AJl * X
ALF \(=0.6366197724 * \operatorname{LOGF}(0.5 * \operatorname{AX})\)
\(Y 0=Y 0+\) ALF * AJO
\(\mathrm{Yl}=\mathrm{Yl} / \mathrm{X}+\mathrm{ALF} * \mathrm{AJl}\)
RETURN
\(\mathrm{X} 3=3.0 / \mathrm{AX}\)
\(\mathrm{FO}=\mathrm{C}(1) \quad \$ \quad \mathrm{THO}=\mathrm{D}(\mathrm{l}) \quad \$ \quad \mathrm{Fl}=\mathrm{G}(\mathrm{l}) \quad \$ \quad \mathrm{TH}=\mathrm{H}(\mathrm{l})\)
DO \(60 \mathrm{M}=2,7\)
\(F 0=F 0\) * \(\mathrm{X} 3+C(M)\)
\(\mathrm{THO}=\mathrm{THO} * \mathrm{X} 3+\mathrm{D}(\mathrm{M})\)
\(\mathrm{Fl}=\mathrm{Fl}{ }^{*} \mathrm{X} 3+\mathrm{G}(\mathrm{M})\)
\(60 \quad \mathrm{THl}=\mathrm{THl} * \mathrm{X} 3+\mathrm{H}(\mathrm{M})\)
\(\mathrm{THO}=\mathrm{THO}+\mathrm{AX} \$ \mathrm{TH}=\mathrm{THl}+\mathrm{AX}\)
\(\mathrm{XS}=1 . / \mathrm{SQRT}(\mathrm{AX})\)
AJO \(=\mathrm{XS}\) * FO * COSF (THO) \(\$ \mathrm{YO}=\mathrm{XS}\) * FO * SINF (THO)
AJl= XS * Fl * COSF(THl) \$ Yl = XS * Fl * SINF (THl)
RETURN
END
C
C PROGRAM FOR CALCULATING THE STARKBROADENING OF HYDROGEN ON THE
C
        BASIS OF THE UNIFIED THEORY WITHOUT LOWER STATE INTERACTION
        DIMENSION DQQ(247), PFAC(6),STRONG(50),FF(1100),SJUU(136,31),
        l SJLL(6,5),RDM(16,3),SLUU(16),SLLL(3),NST(50),DST(50),W(300,5).
        2 DIP(16,3,20),FPAR}(6,50
        COMMON/FDAT/P1,P2,B1,A2,B2,PPFF,CCREAL,CCIMAG
        COMMON/PFW/FIELD(301)
        COMMON/ll/SLOP(16,16,60)
        COMMON/22/DIPOL (16,16,20),NSTEP,FPPP (6,15),NNUU , NNLL, BET
        COMMON/33/AMATR (34,68), BMATR (34,34), CMATR (34,34)
        EQUIVALENCE (SJUU,AMATR)
        FIELD(l) = 0.0
        READ 100, ((W(I,J), J = l,5), I = 1,300)
    100 FORMAT (5E12.4)
    120 READ 150, DEN,TEMP,NNUU,NNLL,GIN,DGG,NTOT,NFAC,(PFAC(I),I=l,6)
    150 FORMAT (2E10.2,I3,I2,2F10.2,2I5/6F10.5)
        IF (EOF,60) 577, 170
C CALCULATION OF MICROFIELD DISTRIBUTION FUNCTION
    170 SHIELD = 0.0898 * DEN**(1./6.)/SQRT (TEMP)
        DO 161 J = l,5
    161 DST(J) = 0.2 * FLOAT (J-l)
        DO l63 I = l,300
        DO 162 J = 1,5
    162 DQQ(J) = W(I,J)
        CALL POLY5 (DST,DQQ,5,STRONG,SLUU,SHIELD,0.1,1)
    163 FIELD(I+l) = SLUU(l)
C CALCULATION OF AVERAGE IONFIELD
        DO lll I = l,300
        AM = I
    lll FF(I) = FIELD(I+l) * AM/l0.
    CALL WEDDLE (0.1,300,FF,BAV,0.)
    DY = l./300. $ Y = 0.
    DO 113 I = 1,10
    Y = Y + DY $ B = 1./Y
    113 FF(I) = WFLD(B) * B**3
    CALL WEDDLE (DY,l0,FF,DB,0.)
    BAV = BAV + DB
C CALCULATION OF WAVELENGTH IN STANDARD AIR
    178 AUU = NNUU
        ALL = NNLL
        SLl = 109678.758 * (1./(ALL*ALL) - l./(AUU*AUU ) )
        SL2 = SLl * SLl
        ALDD = 1.000064328+2949810./(1.46E+10-SL2)+25540./(4.1E+9-SL2)
        ALAM = 1.E+8/(ALDD * SLl)
        PRINT 180, DEN, TEMP, NNUU, NNLL, ALAM, SHIELD
    180 FORMAT (lHl,* DENSITY =* El0.2* TEMPERATURE =*ElO.2,
    l * QUANTUMNUMBERS N UPPER =*I3* N LOWER =*I2* WAVELENGTH =*
    F F8.2* ANGSTROM*/10X,*SHIELDING PARAMETER =*F6.3,10X,
    3 * NO LOWER STATE INTERACTION *//)
        IF (NNUU.LE.l6.AND.NNLL.LE.3) GO TO 200
        PRINT 190
    190 FORMAT (* OVERFLOW OF MATRICES, PROGRAM NOT EXECUTED*)
    577 CALL EXIT
C
C RADIAL MATRIXELEMENTS AND TOTAL LINESTRENGTH
    200 IPI = MOD(NNUU-NNLL,2) + 1
    STOT = 0.
    DO 330 KLL = 1,NNLL
    LLL = KLL - 1 $ ALL = LLL
```

```
    DO 310 KUU = 1,NNUU
    LUU = KUU - l $ AUU = LUU $ RRMM = 0.
    IF (IABS(KUU - KLL).NE.1) GO TO 310
    RRMM = RADMAT(NNUU,LUU,NNLL,LLL) * S3J(ALL,AUU,1.,0.,0.,0.)
    STOT = STOT + (2.*ALL+1.)*(2.*AUU+1.) * RRMM * RRMM
    310 RDM(KUU,KLL) = RRMM
    330 CONTINUE
    SQST = SQRT(STOT)
C
C TRANSFORMATION MATRICES
    DO 370 NKK = l,2
    NN = NNLL
    IF (NKK.EQ.2) NN = NNUU
    ANN = NN $ ANl = 0.5 * (ANN - 1.)
    KQQ = 2 * NN - l $ NRUN = 0
    DO 370 KL = l,NN
    AL = KL - l
    FACC = SQRT(2.*AL+1.)
    IF(NKK.EQ.l) SLLL(KL) = FACC
    IF(NKK.EQ.2) SLUU(KL) = FACC
    DO 370 KM = l,KL
    AM = KM - l $ NRUN = NRUN + l
    QQ = -ANN $ ALIM = ANN - 0.9 - AM
    DO 370 KQ = l,KQQ
    QQ = QQ + l. $ SIIJ = 0. $ ABQQ = ABS (QQ)
    IF (ABQQ.GT.ALIM) GO TO 360
    NQ = ABQQ + O.l
    IF (MOD(NN+NQ+KM+1,2).NE.1) GO TO 360
    AMQ = 0.5 * (AM - QQ) $ APQ = 0.5 * (AM + QQ)
    SIIJ = FACC * S3J(ANl,ANl,AL,AMQ,APQ,-AM)
    360 IF (NKK.EQ.l) SJLL(NRUN,KQ) = SIIJ
    IF (NKK.EQ.2) SJUU(NRUN,KQ) = SIIJ
    370 CONTINUE
C
C DIPOLMATRIXELEMENTS IN PARABOLIC STATES
    NQMAX = NNUU * (NNUU-l) + NNLL * (NNLL-1) + l
    DO 490 MM = l,NQMAX
490 DQQ (MM) = 0.
    MMAX = NNLL + 1 $ KQLL = 2 * NNLL - 1
    DO 500 MM = l,MMAX
    MMU = MM - 1 $ AMU = MMU $ AMLCK = l.
    IF (MM.GT.l) AMLCK = 2.
    NAQU = -NNUU + MMU - 1 $ KQUU = NNUU - MMU
    DO 500 KUU = l,KQUU
    NAQU = NAQU + 2 $ KQU = NNUU + NAQU $ NAQL = -NNLL
    DO 500 KLL = l,KQLL
    NRUN = MMU * KQLL + KLL $ NAQL = NAQL + 1 $ KQL = NNLL + NAQL
    NXDIF = NNUU * NAQU - NNLL * NAQL + 1
    MML = MMU - 2 $ KMINL = NNLL - IABS(NAQL) - l
    DO 470 KAML = l.3
    SIIJ = 0. $ MML = MML + l
    IF (MOD(NNLL+NAQL+MML,2).NE.1) GO TO 470
    IF (IABS(MML).GT.KMINL) GO TO 470
    AML = MML $ AMBL = AMU - AML
    NPHAS = (NNUU + MMU - NAQU + NNLL - MML - NAQL - 2)/2
    PHASE = (-1.)**MOD(NPHAS,2)
    RESULT = 0. $ KBBU = IABS (MMU) + 1 $ KBBL = IABS(MML) + l
    DO 450 KKLU = KBBU,NNUU
    ALU = KKLU - l
    LLU = (KKLU-1) * KKLU / 2 + KBBU
    FACl = SLUU(KKLU) * SJUU(LLU,KQU)
    DO 450 KKLL = KBBL,NNLL
    IF (IABS(KKLU - KKLL).NE.l) GO TO 450
```

```
    ALL. = KKLL - I
    LLL = (KKLL - l) * KKLL / 2 + KBBL
    FAC2 = SLLL(KKLL) * SJLL(LLL,KQL)
    FACC = FAC2 * FACl * S3J(ALL,ALU,1., AML, -AMU,AMBL)
    RESULT = RESULT + FACC * RDM(KKLU,KKLL)
    4 5 0
    CONTINUE
    SIIJ = RESULT * PHASE / SQST
    PRINT 425,NAQU,MMU,NAQL,MML,SIIJ
4 2 5 ~ F O R M A T ~ ( 1 6 X , * ~ ( * , 2 I 3 , * ~ I ~ D ~ I * , 2 I 3 , * ~ ) ~ = * E 1 6 . 8 )
    IF (NXDIF.GT.O) DQQ(NXDIF) = DQQ(NXDIF) + AMLCK*RESULT*RESULT/STOT
4 7 0 ~ D I P ( K U U , K A M L , N R U N ) ~ = ~ S I I J ~
5 0 0 ~ C O N T I N U E ~
    PRINT }39
393 FORMAT(/)
    DO 440 MM= 1,MMAX
    MMU = MM - l $ KQUU = NNUU - MMU $ AMLCK = l
    IF (MMU.GT.O) AMLCK = 2.
    DO 440 KLL = 1,KQLL
    NRUN = MMU * KQLL + KLL
    DO 440 NRA = l,KQUU
    DO 440 NRB = l,KQUU
    SIIJ = 0.
    DO 420 KAML = l,3
4 2 0 ~ S I I J ~ = ~ S I I J ~ + ~ A M L C K * D I P ( N R A , K A M L , N R U N ) * D I P ( N R B , K A M L , N R U N )
4 4 0 ~ D I P O L ( N R A , N R B , N R U N ) ~ = ~ S I I J ~
    DST(l) = DQQ(l) $ NST(l) = 0 $ KMST = IPI - l
    DO 430 NXDIF = 2,NQMAX
    IF (DQQ(NXDIF).LT.1.E-20) GO TO 430
    KMST = KMST + l $ DST(KMST) = DQQ(NXDIF) $ NST(KMST) = NXDIF-l
430 CONTINUE
    PRINT 355, (NST(I),DST(I),I = 1,KMST)
355 FORMAT (16X,*X =*I3,F20.10)
    PRINT }39
C
C K OPERATOR MATRIX
    NTIME = KLOCK(0) $ NCMl = NNUU - l
    DO 600 MM = l,MMAX
    MMU = MM - l $ AMU = MMU
    DO 600 NRC = l,NCMl
    NRUN = MMU * NCM1 + NRC $ KCQU = NNUU + NRC
    NAQU = -NNUU + MMU - 1 $ KQUU = NNUU - MMU
    DO 600 NRA = 1,KQUU
    NAQU = NAQU + 2 $ KAQU = NNUU + NAQU $ NBQU = -NNUU + MMU - l
    DO }600\mathrm{ NRB = 1,NRA
    NBQU = NBQU + 2 $ KBQU = NNUU + NBQU $ SIIJ = 0.
    DO 540 JLAU = MM,NNUU
    ALU = JLAU - I
    LALU = (JLAU-l)*JLAU/2 + MM
    FACl = SJUU(LALU,KAQU) * SJUU(LALU,KBQU)
    FAC2 = 0. $ KMINL = MINO(JLAU,NNUU-NRC)
    DO 520 JMCU = 1,KMINL
    IF (MOD(NNUU+NRC+JMCU,2).EQ.1) GO TO 520
    LCLU = LALU - MM + JMCU $ AMC = 1.
    IF (JMCU.GT.l) AMC = 2.
    NPHAS = MM + JMCU + NRC - (NAQU+NBQU)/2
    PHASE = (-1.)**MOD(NPHAS, 2)
    FAC2 = FAC2 + PHASE*AMC* (SJUU (LCLU,KCQU) )**2
5 2 0 ~ C O N T I N U E ~
540 SIIJ = SIIJ + FACl * FAC2/(2.*ALU+l.)
    SLOP(NRA,NRB,NRUN) = SIIJ
600 SLOP(NRB,NRA,NRUN) = SIIJ
    NTIME = -NTIME + KLOCK(0)
    PRINT 617, NTIME
```

    BASIC CONSTANTS AND ARRAY FOR G-FUNCTION CONSTANTS
    SDEN = SQRT(DEN)
    FAC \(=2064.936\) * TEMP * SQRT (TEMP/DEN)
    \(\mathrm{CFAC}=4.5645 \mathrm{E}-7\) * SDEN \(/\) TEMP
    DEBROG \(=2.1027 \mathrm{E}-6 /\) SQRT (TEMP)
    ANN \(=\) NNUU
    RMIN \(=\) DEBROG + ANN * ANN * 7.9376E-9
    BET \(=5.6558 \mathrm{E}-5 *\) DEN \(^{*} *(1 . / 6\).
    \(\mathrm{ASY}=0.0\)
    DO 270 NRC = IPI,KMST
    \(\mathrm{QC}=\mathrm{NST}(\mathrm{NRC}) \quad \$ \quad \mathrm{C}=\mathrm{CFAC} * \mathrm{QC}\)
    Pl \(=-1.671086\) * \(\mathrm{FAC} * \mathrm{C}\) * SQRT(C)
    \(\mathrm{BS}=3 .{ }^{*} \mathrm{QC}{ }^{*}\) DEBROG / RMIN
    STRONG(NRC) \(=0.269-2 . *(((1 .-\operatorname{COSF}(B S)) / B S+S I N F(B S)) / B S-C O S I N T(B S))\)
    PPFF \(=-1.128379\) * FAC * C * C
    P2 \(=\) PPFF * (STRONG (NRC) -2. *LOGF (2. * \({ }^{*}\) ) )
    \(\operatorname{FPAR}(1, N R C)=P 1 \$ \operatorname{FPAR}(2, N R C)=P 2\)
    \(\operatorname{FPAR}(3, \operatorname{NRC})=0.5 *(\mathrm{P} 2 / \mathrm{Pl}) * * 2 \$ \mathrm{FIN}=\operatorname{LOGF}(\mathrm{QC})\)
    \(\mathrm{A} 2=\mathrm{P} 2 *((\operatorname{PFAC}(3) * F I N+\operatorname{PFAC}(2)) * F I N+\operatorname{PFAC}(1))\)
    \(B 2=(\operatorname{PFAC}(6) * F I N+\operatorname{PFAC}(5)) * F I N+\operatorname{PFAC}(4)\)
    IF (B2.LT.0.) A2 \(=0\).
    \(\operatorname{FPAR}(4, \mathrm{NRC})=\mathrm{A} 2 \$ \operatorname{FPAR}(5, \mathrm{NRC})=\mathrm{B} 2 \$ \operatorname{FPAR}(6, \mathrm{NRC})=\operatorname{PPFF}\)
    270
ASY $=$ ASY $+2 .{ }^{*}$ Pl ${ }^{*} \operatorname{DST}(N R C)$
DO $275 \mathrm{NRC}=1$, NCMI
NCC $=$ NRC * NNUU
DO 278 NRA = IPI,KMST
IF (NST (NRA) .EQ.NCC) GO TO 279
278 CONTINUE
$279 \operatorname{FPPP}(1, N R C)=\operatorname{FPAR}(1, N R A) \$ \operatorname{FPPP}(2, N R C)=\operatorname{FPAR}(2, N R A)$
$\operatorname{FPPP}(3, N R C)=\operatorname{FPAR}(3, N R A) \$ \operatorname{FPPP}(4, \operatorname{NRC})=\operatorname{FPAR}(4, N R A)$
$\operatorname{FPPP}(5, \operatorname{NRC})=\operatorname{FPAR}(5, N R A)$
$\operatorname{FPPP}(6, \operatorname{NRC})=\operatorname{FPAR}(6, N R A)$
PRINT 220
220 FORMAT (/13X*P1*18X*P2*18X*B1*18X*A2*18X*B2*17X*STRONG*/)
PRINT 240, ((FPAR(K,I), $K=1,5), S T R O N G(I), I=I P I, K M S T)$
240 FORMAT (6E20.4)
PRINT 280, FAC, CFAC, BET, ASY, DEBROG, BAV, NFAC
280 FORMAT(/* FAC $={ }^{*}$ Ell.4,* CFAC $=* E l 1.4$,* BET $={ }^{*}$ Ell.4,* ASY $={ }^{*}$ Ell. 4 ,
* DEBROG $=$ *Ell.4,* BAV $=$ *F7.4,* INTEGRATIONFACTOR $=$ *I2//
5X*DOM*8X*DLAM*8X*ITOT*6X*IHOLTS*8X*ASY*10X*WING*7X*WHOLTS*7X,
*WWOO*8X*WWBB*8X*WWNG*8X*TIME STEPS*/)
ADLFAC $=4.23538 \mathrm{E}-15 *$ SDEN $*$ ALAM * ALAM
c
C CALCULATION OF THE IONFIELD INTEGRAL
$\mathrm{N} 12=12 * \mathrm{NFAC} \$ \mathrm{AN12}=\mathrm{N} 12 \$ \$ \mathrm{~N} 30=30 * \mathrm{NFAC} \$$ AN3O $=\mathrm{N} 30$
DMCRT $=$ BET * BAV * NST (KMST) * 5 .
$\mathrm{G}=\mathrm{GIN}-\mathrm{DGG}$
DO $950 \mathrm{MM}=1, \mathrm{NTOT}$
NTIME $=\operatorname{KLOCK}(0)$
NSTEP $=0 \quad \$ \quad G=G+D G G$
DOM $=10$. ** G $\$$ DLAM $=$ ADLFAC * DOM
WING52 $=-0.2992067103 *$ ASY $/(S Q R T(D O M) *$ DOM * DOM)
FHOLTS $=0 . \quad \$$ AWING $=0$.
DO 815 NRC = IPI,KMST
$\mathrm{Pl}=\operatorname{FPAR}(1, \mathrm{NRC}) \quad \$ \quad \operatorname{P2}=\operatorname{FPAR}(2, \mathrm{NRC}) \quad \$ \quad \mathrm{Bl}=\operatorname{FPAR}(3, \mathrm{NRC})$
$\mathrm{A} 2=\operatorname{FPAR}(4, \mathrm{NRC}) \quad \$ \quad \mathrm{~B} 2=\operatorname{FPAR}(5, \mathrm{NRC}) \quad \$ \quad \operatorname{PPFF}=\operatorname{FPAR}(6, \mathrm{NRC})$
CALL FOUTR(DOM)

$\mathrm{QC}=\mathrm{NST}(\mathrm{NRC}) \$ \mathrm{BETFAC}=\mathrm{BET} * \mathrm{QC} \$ \mathrm{BCRIT}=\mathrm{DOM} / \mathrm{BETFAC}$
815 FHOLTS $=$ FHOLTS + DST $($ NRC $) ~ * ~ W F L D ~(B C R I T) ~ / ~ B E T F A C ~$
IF (DOM.GT.DMCRT) GO TO 985

```
    AIRES = 0
    IF(DOM.GT.(-3.*P2)) GO TO 840
    ANQ1 = NST(KMST)
    BCRIT = (DOM - P2)/(ANQ1*BET)
    DB = BCRIT/AN12 $ B = 0.
    DO 820 J = l,N12
    B = B + DB
    LD (B
    CALL WEDDLE (DB,N12,FF,AIII,0.)
    AIRES = AIII
    DY = 1./(BCRIT*AN30) $ Y = 0.
    DO 830 J = l,N30
    Y = Y + DY $ B = l./Y
830 FF(J) = B * B * AIIM(DOM,B) * WFLD(B)
    CALL WEDDLE (DY,N30,FF,AIII,0.)
    AIRES = AIRES + AIII
    GO TO 980
840 BCRCR = DOM/BET
    EPSPS = -P2/BET
    DO 957 NQ = IPI,KMST
    ANQ = NST(NQ)
    BCR = BCRCR/ANQ $ EPS = EPSPS/ANQ
    IF (NQ.EQ.IPI) GO TO 907
    SLl = 1./(GAM - BCR)
    G0 T0 908
907
908 SL2 = 1./EPS
    SL3 = 1./(BCR + EPS ) $ SL4 = 1./(BCR - EPS )
    IF (NQ.EQ.KMST) GO TO 911
    ANQ1 = NST(NQ+l)
    GAM = 0.5 * (BCR-EPS+(BCRCR+EPSPS)/ANQ1)
    G0 T0 912
911 GAM = 0.5 * (BCR-EPS)
912 SL5 = 1./(BCR-GAM)
    CRIT = SL2 - SL5
904 Y = SLl
    IF (NQ.EQ.IPI) GO TO 913
    B = BCR + 1./Y
    FA = AIIM(DOM,B) * WFLD(B)/(Y * Y)
    GO TO 9l4
913 FA = 0.
914 DY = (SL2 - SLl)/AN12
    DO 917 J = l,Nl2
    Y = Y + DY $ Yl = l./Y $ B = BCR + Yl
917 FF(J) = Yl * Yl * AIIM(DOM,B) * WFLD(B)
    CALL WEDDLE (DY,N12,FF,AIII,FA)
    AIRES = AIRES + AIII
    Y = SL3 $ B = 1./Y
    FA = B * B * AIIM(DOM,B) * WFLD(B)
    DY = (SL4 - SL3)/AN12
    DO 927 J = l,N12
    Y = Y + DY $ B = l./Y
927 FF(J) = B * B * AIIM(DOM,B) * WFLD(B)
    CALL WEDDLE (DY,N12,FF,AIII,FA)
    AIRES = AIRES + AIII
    IF(CRIT.LE.O.) GO TO 977
    Y = SL5 $ B = BCR - 1./Y
    FA = AIIM(DOM,B) * WFLD(B)/(Y * Y)
    DY = CRIT/ANI2
    DO 937 J = l,N12
    Y = Y + DY $ Yl = l./Y $ B = BCR - Yl
    FF(J) = Yl * Yl * AIIM(DOM,B) * WFLD(B)
    CALL WEDDLE (DY,N12,FF,AIII,FA)
```

```
    AIRES = AIRES + AIII
    957 CONTINUE
    IF(GAM.LT. 5.) GO TO 968
    Y = l./GAM $ DY = (0.2 - Y)/ANl2
    FA = GAM * GAM * AIIM(DOM,GAM) * WFLD(GAM)
    DO 967 J = l,Nl2
    Y = Y + DY $ B = 1./Y
    967 FF(J) = B * B * AIIM(DOM,B) * WFLD(B)
    CALL WEDDLE (DY,Nl2,FF,AIII,FA)
    AIRES = AIRES + AIII
    SL4 = 0.2
    GO TO 977
    968 SL4 = l./GAM
    977 B = 0.
    DB = 1./(SL4 * AN30)
    DO 947 J = l,N30
    B = B + DB
    947 FF(J) = AIIM(DOM,B) * WFLD(B)
    CALL WEDDLE (DB,N30,FF,AIII,O.)
    AIRES = AIRES + AIII
    980 WWBB = (AIIM(DOM,BAV) + FHOL'TS)/WING52
    G0 T0 990
985 AIRES = AIIM(DOM,BAV) + FHOLTS
    WWBB = 0.
    990 WING = AIRES/WING52
    WINHOL = FHOLTS/WING52
    WWOO = (AIIM(DOM, 0.) + FHOLTS)/WING52
    WWNG = (AWING + FHOLTS)/WING52
    NTIME = -NTIME + KLOCK(O)
    950 PRINT 978,DOM,DLAM,AIRES,FHOLTS,WING52,WING,WINHOL,WWOO,WWBB,
    l WWNG,NTIME,NSTEP
    978 FORMAT (10E12.4,I10,I6)
    GO TO 120
    END
C
C
    FUNCTION AIIM(DOM,B)
C CALCULATION OF I(DOM,B) FOR THE CASE OF NO LOWER STATE INTERACTION
    DIMENSION DRRF(15), DIIF(15), FMATR(15,15)
    COMMON/FDAT/P1,P2,B1,A2,B2,PPFF,CCREAL,CCIMAG
    COMMON/ll/SLOP(16,16,60)
    COMMON/22/DIPOL (16,16,20),NSTEP,FPPPP (6,15),NNUU ,NNLL, BET
    COMMON/33/AMATR(34,68), BMATR(34,34), CMATR(34,34)
    AIIM = 0. $ NSTEP = NSTEP + l
    MMAX = NNLL + l $ KQLL = 2 * NNLL - l $ NCMl = NNUU - 1
    DO 800 MM = l,MMAX
    MMU = MM - l $ NUPP = NNUU - MMU $ NQL = -NNLL'
    DO 800 KLL = 1,KQLL
    NQL = NQL + I
    NRUN = MMU * KQLL + KLL $ NBQU = -NNUU + MMU - l
    DO 750 NRB = l,NUPP
    NBQU = NBQU + 2 $ AXDIF = NNUU * NBQU - NNLL * NQL
    DOMRB = DOM - BET * B * AXDIF
    DO 22O NRC = l,NCMI
    Pl = FPPP(1,NRC) $ P2 = FPPP(2,NRC) $ Bl = FPPP(3,NRC)
    A2 = FPPP(4,NRC) $ B2 = FPPP(5,NRC) $ PPFF = FPPP(6,NRC)
    CALL FOUTR(DOMRB)
    DRRF(NRC) = CCREAL
220 DIIF (NRC) = CCIMAG
    DO 700 NRA = l,NUPP
    ARRR = 0. $ AIII = 0.
    DO 600 NRC = l,NCMI
```

```
        MQC = MMU * NCMl + NRC
            TEMP = SLOP(NRB,NRA,MQC)
            ARRR = ARRR + TEMP * DRRF (NRC)
    6 0 0
    AIII = AIII + TEMP * DIIF(NRC)
    AIII = AIII * 6.2831853072
    AMATR(NRB,NRA) = AIII $ CMATR(NRB,NRA) = AIII
    700 BMATR(NRB,NRA) = ARRR * 6.2831853072
750 BMATR(NRB,NRB) = BMATR(NRB,NRB) + DOMRB
    CALL ENVERS (NUPP)
    DO 320 NRB = l,NUPP
    DO 320 NRA = l,NUPP
    TEMP = 0
    DO 300 NRC = l,NUPP
    300 TEMP = TEMP + BMATR(NRB,NRC) * AMATR(NRC,NRA+NUPP)
    320 FMATR(NRB,NRA) = TEMP
    DO 420 NRB = l,NUPP
    DO 420 NRA = l,NUPP
    TEMP = 0.
    DO 400 NRC = l,NUPP
    TEMP = TEMP + FMATR(NRB,NRC) * BMATR(NRC,NRA)
420 AMATR(NRB,NRA) = TEMP + CMATR(NRB,NRA)
    CALL ENVERS (NUPP)
    DO 795 NRB = l,NUPP
    DO 795 NRA = l,NUPP
795 AIIM = AIIM + DIPOL(NRB,NRA,NRUN)*AMATR(NRB,NRA+NUPP)
800 CONTINUE
    AIIM = AIIM * 0.3183099 $ RETURN
    END
C
C
```

LOW FREQUENCY COMPONENT OF THE ELECTRIC MICROFIELDS

| $\tau_{0} / \mathrm{D}=0$ | $=0.2$ | $=0.4$ |
| :---: | :---: | :---: |
| $4.2200-003$ | $6.9547-003$ | $1.1584-002$ |
| $1.6670-002$ | $2.7217-002$ | $4.4735-002$ |
| $3.6640-002$ | $5.9060-002$ | $9.5010-002$ |
| $6.3080-002$ | $9.9852-002$ | $1.5606-001$ |
| $9.4600-002$ | $1.4638-001$ | $2.2090-001$ |
| $1.2959-001$ | $1.9523-001$ | $2.8304-001$ |
| $1.6636-001$ | $2.4312-001$ | $3.3742-001$ |
| $2.0323-001$ | $2.8721-001$ | $3.8073-001$ |
| $2.3864-001$ | $3.2532-001$ | $4.1151-001$ |
| $2.7122-001$ | $3.5597-001$ | $4.2978-001$ |
| $2.9987-001$ | $3.7845-001$ | $4.3663-001$ |
| $3.2378-001$ | $3.9268-001$ | $4.3381-001$ |
| $3.4246-001$ | $3.9910-001$ | $4.2331-001$ |
| $3.5570-001$ | $3.9855-001$ | $4.0710-001$ |
| $3.6357-001$ | $3.9206-001$ | $3.8700-001$ |
| $3.6633-001$ | $3.8077-001$ | $3.6450-001$ |
| $3.6445-001$ | $3.6584-001$ | $3.4084-001$ |
| $3.5850-001$ | $3.4831-001$ | $3.1694-001$ |
| $3.4911-001$ | $3.2911-001$ | $2.9347-001$ |
| $3.3694-001$ | $3.0904-001$ | $2.7090-001$ |
| $3.2265-001$ | $2.8872-001$ | $2.4953-001$ |
| $3.0684-001$ | $2.6865-001$ | $2.2952-001$ |
| $2.9005-001$ | $2.4918-001$ | $2.1096-001$ |
| $2.7275-001$ | $2.3058-001$ | $1.9385-001$ |
| $2.5537-001$ | $2.1300-001$ | $1.7816-001$ |

LOW FREQUENCY COMPONENT OF THE ELECTRIC MICROFIELDS

| $\mathrm{r}_{\mathrm{d}} / \mathrm{D}=0$ | $=0.2$ |
| :---: | :---: |
| 2.3822-001 | 1.9654-001 |
| 2.2156-001 | 1.8124-001 |
| 2.0557-001 | 1.6709-001 |
| 1.9037-001 | 1.5406-001 |
| 1.7606-001 | 1.4211-001 |
| 1.6268-001 | 1.3116-001 |
| 1.5024-001 | 1.2117-001 |
| 1.3873-001 | 1.1204-001 |
| 1.2812-001 | 1.0371-001 |
| 1.1837-001 | 9.6123-002 |
| 1.0942-001 | 8.9201-002 |
| 1.0124-001 | 8.2886-002 |
| 9.3751-002 | 7.7122-002 |
| 8.6913-002 | 7.1857-002 |
| 8.0670-002 | 6.7044-002 |
| 7.4969-002 | 6.2639-002 |
| 6.9762-002 | 5.8603-002 |
| 6.5004-002 | 5.4901-002 |
| 6.0652-002 | 5.1502-002 |
| 5.6670-002 | 4.8376-002 |
| 5.3023-002 | 4.5498-002 |
| 4.9678-002 | 4.2844-002 |
| 4.6606-002 | 4.0395-002 |
| 4.3781-002 | 3.8130-002 |
| 4.1180-002 | 3.6034-002 |
| 3.8783-002 | 3.4092-002 |
| 3.6571-002 | 3.2289-002 |
| 3.4529-002 | 3.0613-002 |
| 3.2640-002 | 2.9054-002 |
| 3.0890-002 | 2.7601-002 |
| $2.9266-002$ | 2.6247-002 |
| 2.7756-002 | 2.4981-002 |
| 2.6347-002 | 2.3798-002 |
| 2.5030-002 | 2.2690-002 |
| 2.3800-002 | 2.1652-002 |
| 2.2655-002 | 2.0678-002 |
| 2.1588-002 | 1.9764-002 |
| 2.0592-002 | 1.8904-002 |
| 1.9662-002 | 1.8095-002 |
| 1.8790-002 | 1.7332-002 |
| 1.7969-002 | 1.6614-002 |
| 1.7197-002 | 1.5935-002 |
| 1.6470-002 | 1.5295-002 |
| 1.5786-002 | 1.4689-002 |
| 1.5140-002 | 1.4116-002 |
| 1.4531-002 | 1.3573-002 |
| 1.3956-002 | 1.3059-002 |
| 1.3413-002 | 1.2571-002 |
| 1.2898-002 | 1.2108-002 |
| 1.2410-002 | 1.1669-002 |
| 1.1947-002 | 1.1251-002 |
| 1.1508-002 | 1.0854-002 |
| 1.1092-002 | 1.0477-002 |
| 1.0696-002 | 1.0117-002 |
| 1.0320-002 | 9.7745-003 |

2.3822-001 2.2156-001 2.0557-001 1.9037-001 1.7606-001
1.6268-001 1.5024-001 1.3873-001 1.2812-001 1.0942-001 1.0124-001 -002 8.6913-002 8.0670-002
7.4969-002 -002 6.0652-002 5.6670-002 5.3023-002 4.9678-002 4.6606-002 4.3781-002 4.1180-002 3.8783-002 3.6571-002 3.4529-002 3.2640-002 2.9266-002 2.7756-002 2.6347-002 2.5030-002 2.3800-002 2.2655-002 $2.1588-002$ 1.9662-002 1.8790-002
1.7969-002 1.7197-002 1.6470-002 1.5786-002 $1.5140-002$ 1.4531-002 1.3956-002 1.3413-002 1.2898-002 -002 1.1508 1.069 1.0320-002
$=0.4$
1.6382-00
1.5075-001
1.3885-001
1.2804-001
1.1821-001
1.0928-001
1.0117-001 9.3793-002 8.7081-002 8.0967-002
7.5392-002 7.0303-002 6.5651-002
6.1393-002
5.7491-002
5.3910-002
5.0620-002
4.7591-002
4.4801-002
4.2226-002
3.9847-002
3.7647-002
3.5608-002
3.3716-002
3.1959-002
3.0325-002
2.8803-002
2.7383-002
2.6058-002
2.4819-002
2.3659-002
2.2573-002
2.1554-002
2.0597-002
1.9697-002
1.8850-002
1.8053-002
1.7302-002
1.6593-002
1.5923-002
1.5290-002
1.4692-002
1.4125-002
1.3588-002
1.3079-002
1.2596-002
1.2138-002
1.1702-002
1.1288-002
1.0894-002
1.0519-002
1.0161-002
9.8208-003
9.4959-003
9.1858-003
$=0.8$
$\begin{array}{ll}1.2077-001 & 26 \\ 1.1129-001 & 27 \\ 1.0273-001 & 28 \\ 9.5010-002 & 29 \\ 8.8028-002 & 30 \\ 8.1696-002 & 31 \\ 7.5951-002 & 32 \\ 7.0731-002 & 33 \\ 6.5969--002 & 34 \\ 6.1626-002 & 35 \\ 5.7662-002 & 36 \\ 5.4030-002 & 37 \\ 5.0700-002 & 38 \\ 4.7650-002 & 39 \\ 4.4839-002 & 40 \\ 4.2246-002 & 41 \\ 3.9859-002 & 42 \\ 3.7647-002 & 43 \\ 3.5594-002 & 44 \\ 3.3696-002 & 45\end{array}$
3.1932-002 46
3.0289-002 47
$2.8766-00248$
2.7348-002 49
2.6021-002 50
2.4785-002 51
2.3631-002 52
2.2543-002 53
2.1527-002 54
2.0578-002 55
$1.9681-00256$
1.8839-002 57
$1.8052-00258$
$1.7302-00259$
1.6593-002
1.5928-002
1.5294-002
1.4693-002
1.4129-002
1.3592-002
1.3079-002
1.2599-002
1.2142-002
1.1704-002
1.1294-002
1.0903-002
1.0525-002
1.0168-002
9.8296-003
9.5003-003
9.1900-003
8.8965-003
8.6096-003
8.3371-003
8.0805-003

LOW FREQUENCY COMPONENT OF THE ELECTRIC MICROFIELDS

| $\mathrm{ra}_{3} / \mathrm{D}=0$ | $=0.2$ | $=0.4$ | $=0.6$ | $=0.8$ | $10 \beta$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $9.9640-003$ | 9.4481-003 | 8.8896-003 | 8.3461-003 | 7.8286-003 | 81 |
| 9.6250-003 | 9.1370-003 | 8.6066-003 | 8.0876-003 | 7.5880-003 | 82 |
| 9.3020-003 | 8.8401-003 | 8.3360-003 | 7.8402-003 | 7.3633-003 | 83 |
| 8.9940-003 | 8.5568-003 | 8.0772-003 | 7.6032-003 | 7.1431-003 | 84 |
| 8.7000-003 | 8.2863-003 | 7.8295-003 | 7.3763-003 | 6.9313-003 | 85 |
| 8.4180-003 | 8.0279-003 | 7.5925-003 | 7.1588-003 | 6.7349-003 | 86 |
| 8.1490-003 | 7.7809-003 | 7.3654-003 | 6.9504-003 | 6.5428-003 | 87 |
| 7.8920-003 | 7.5447-003 | 7.1480-003 | 6.7505-003 | 6.3557-003 | 88 |
| 7.6450-003 | 7.3188-003 | 6.9395-003 | 6.5589-003 | 6.1826-003 | 89 |
| 7.4100-003 | 7.1025-003 | 6.7397-003 | 6.3749-003 | 6.0139-003 | 90 |
| 7.1850-003 | 6.8955-003 | 6.5481-003 | 6.1982-003 | 5.8473-003 | 91 |
| 6.9700-003 | 6.6971-003 | 6.3641-003 | 6.0284-003 | 5.6932-003 | 92 |
| 6.7640-003 | 6.5070-003 | 6.1876-003 | 5.8652-003 | 5.5444-003 | 93 |
| 6.5680-003 | 6.3246-003 | 6.0179-003 | 5.7081-003 | 5.3958-003 | 94 |
| 6.3800-003 | 6.1496-003 | 5.8548-003 | 5.5569-003 | 5.2582-003 | 95 |
| 6.2000-003 | 5.9814-003 | 5.6979-003 | 5.4113-003 | 5.1267-003 | 96 |
| 6.0270-003 | 5.8199-003 | 5.5469-003 | 5.2708-003 | 4.9939-003 | 97 |
| 5.8600-003 | 5.6645-003 | 5.4014-003 | 5.1353-003 | 4.8699-003 | 98 |
| 5.6980-003 | 5.5149-003 | 5.2611-003 | 5.0046-003 | 4.7530-003 | 99 |
| 5.5400-003 | 5.3709-003 | 5.1258-003 | 4.8783-003 | 4.6335-003 | 100 |
| 5.4133-003 | 5.2320-003 | 4.9952-003 | 4.7563-003 | 4.5206-003 | 101 |
| 5.2692-003 | 5.0980-003 | 4.8690-003 | 4.6382-003 | 4.4154-003 | 102 |
| 5.1306-003 | 4.9686-003 | 4.7470-003 | 4.5240-003 | 4.3074-003 | 103 |
| 4.9970-003 | 4.8435-003 | 4.6290-003 | 4.4133-003 | 4.2035-003 | 104 |
| 4.8682-003 | 4.7227-003 | 4.5148-003 | 4.3062-003 | 4.1081-003 | 105 |
| 4.7441-003 | 4.6057-003 | 4.4043-003 | 4.2024-003 | 4.0101-003 | 106 |
| 4.6244-003 | 4.4925-003 | 4.2973-003 | 4.1017-003 | 3.9140-003 | 107 |
| 4.5090-003 | 4.3828-003 | 4.1936-003 | 4.0042-003 | 3.8269-003 | 108 |
| 4.3976-003 | 4.2765-003 | 4.0932-003 | 3.9097-003 | 3.7380-003 | 109 |
| 4.2900-003 | 4.1736-003 | 3.9960-003 | 3.8181-003 | 3.6490-003 | 110 |
| 4.1861-003 | 4.0737-003 | 3.9018-003 | 3.7294-003 | 3.5690-003 | 111 |
| 4.0857-003 | 3.9770-003 | 3.8106-003 | 3.6435-003 | 3.4886-003 | 112 |
| 3.9887-003 | 3.8832-003 | 3.7224-003 | 3.5603-003 | 3.4065-003 | 113 |
| 3.8948-003 | 3.7923-003 | 3.6370-003 | 3.4800-003 | 3.3331-003 | 114 |
| 3.8041-003 | 3.7041-003 | 3.5545-003 | 3.4024-003 | 3.2610-003 | 115 |
| 3.7164-003 | 3.6188-003 | $3.4747-003$ | 3.3274-003 | 3.1860-003 | 116 |
| 3.6314-003 | 3.5361-003 | 3.3977-003 | 3.2551-003 | 3.1188-003 | 117 |
| 3.5492-003 | 3.4560-003 | 3.3233-003 | 3.1854-003 | 3.0549-003 | 118 |
| 3.4696-003 | 3.3785-003 | 3.2515-003 | 3.1182-003 | 2.9870-003 | 119 |
| 3.3925-003 | $3.3034-003$ | 3.1822-003 | 3.0534-003 | 2.9259-003 | 120 |
| 3.3177-003 | 3.2307-003 | 3.1153-003 | 2.9911-003 | 2.8694-003 | 121 |
| 3.2453-003 | 3.1603-003 | 3.0507-003 | 2.9309-003 | 2.8086-003 | 122 |
| 3.1751-003 | 3.0922-003 | 2.9883-003 | 2.8728-003 | 2.7527-003 | 123 |
| $3.1070-003$ | 3.0262-003 | $2.9279-003$ | 2.8167-003 | 2.7027-003 | 124 |
| 3.0410-003 | 2.9623-003 | 2.8694-003 | 2.7623-003 | 2.6481-003 | 125 |
| 2.9769-003 | 2.9003-003 | 2.8127-003 | $2.7096-003$ | 2.5961-003 | 126 |
| 2.9147-003 | 2.8402-003 | 2.7576-003 | 2.6583-003 | 2.5509-003 | 127 |
| 2.8543-003 | 2.7819-003 | $2.7040-003$ | 2.6083-003 | 2.5013-003 | 128 |
| 2.7957-003 | 2.7253-003 | 2.6519-003 | 2.5594-003 | 2.4521-003 | 129 |
| 2.7387-003 | 2.6702-003 | 2.6009-003 | 2.5116-003 | $2.4100-003$ | 130 |
| 2.6834-003 | 2.6167-003 | 2.5512-003 | $2.4646-003$ | $2.3643-003$ | 131 |
| 2.6296-003 | 2.5647-003 | $2.5025-003$ | 2.4185-003 | 2.3170-003 | 132 |
| 2.5774-003 | 2.5141-003 | 2.4549-003 | $2.3732-003$ | 2.2769-003 | 133 |
| $2.5266-003$ | 2.4648-003 | 2.4082-003 | 2.3286-003 | 2.2346-003 | 134 |
| $2.4772-003$ | 2.4169-003 | $2.3625-003$ | $2.2848-003$ | 2.1890-003 | 135 |

LOW FREQUENCY COMPONENT OF THE ELECTRIC MICROFIELDS

| $\mathrm{r}_{0} / \mathrm{D}=0$ | $=0.2$ | $=0.4$ |
| :---: | :---: | :---: |
| 2.4291-003 | 2.3703-003 | 2.3178-003 |
| 2.3824-003 | 2.3249-003 | 2.2740-003 |
| 2.3369-003 | 2.2807-003 | 2.2311-003 |
| 2.2926-003 | 2.2378-003 | 2.1892-003 |
| 2.2495-003 | 2.1961-003 | 2.1482-003 |
| 2.2075-003 | 2.1555-003 | 2.1083-003 |
| 2.1667-003 | 2.1162-003 | 2.0694-003 |
| 2.1268-003 | 2.0779-003 | 2.0314-003 |
| 2.0880-003 | 2.0408-003 | 1.9945-003 |
| 2.0502-003 | 2.0048-003 | 1.9587-003 |
| 2.0134-003 | 1.9698-003 | 1.9238-003 |
| 1.9775-003 | 1.9358-003 | 1.8899-003 |
| 1.9424-003 | 1.9029-003 | 1.8571-003 |
| 1.9083-003 | 1.8708-003 | 1.8251-003 |
| 1.8750-003 | $1.8397-003$ | 1.7941-003 |
| 1.8425-003 | 1.8094-003 | 1.7640-003 |
| 1.8108-003 | 1.7799-003 | 1.7347-003 |
| 1.7798-003 | 1.7511-003 | 1.7062-003 |
| 1.7496-003 | 1.7230-003 | 1.6785-003 |
| 1.7201-003 | 1.6955-003 | 1.6515-003 |
| 1.6913-003 | 1.6687-003 | 1.6252-003 |
| 1.6632-003 | 1.6424-003 | 1.5995-003 |
| 1.6357-003 | 1.6166-003 | 1.5744-003 |
| 1.6089-003 | 1.5913-003 | 1.5498-003 |
| 1.5826-003 | 1.5665-003 | 1.5257-003 |
| 1.5570-003 | 1.5421-003 | 1.5021-003 |
| 1.5320-003 | 1.5181-003 | 1.4789-003 |
| 1.5075-003 | 1.4945-003 | 1.4562-003 |
| 1.4836-003 | 1.4713-003 | 1.4338-003 |
| 1.4602-003 | 1.4485-003 | 1.4118-003 |
| 1.4373-003 | 1.4260-003 | 1.3903-003 |
| 1.4149-003 | 1.4039-003 | 1.3691-003 |
| 1.3930-003 | 1.3822-003 | 1.3482-003 |
| 1.3716-003 | 1.3608-003 | 1.3278-003 |
| 1.3506-003 | 1.3398-003 | 1.3078-003 |
| 1.3301-003 | 1.3192-003 | 1.2881-003 |
| 1.3100-003 | 1.2990-003 | 1.2688-003 |
| 1.2904-003 | 1.2791-003 | 1.2499-003 |
| 1.2711-003 | 1.2597-003 | 1.2314-003 |
| 1.2523-003 | 1.2406-003 | 1.2132-003 |
| 1.2339-003 | 1.2219-003 | 1.1953-003 |
| 1.2158-003 | 1.2036-003 | 1.1778-003 |
| 1.1981-003 | 1.1857-003 | 1.1607-003 |
| 1.1808-003 | 1.1681-003 | 1.1438-003 |
| $1.1638-003$ | 1.1510-003 | 1.1273-003 |
| 1.1471-003 | 1.1342-003 | 1.1111-003 |
| 1.1308-003 | 1.1178-003 | 1.0952-003 |
| 1.1148-003 | 1.1017-003 | 1.0796-003 |
| 1.0992-003 | 1.0860-003 | 1.0643-003 |
| 1.0838-003 | 1.0707-003 | 1.0494-003 |
| 1.0687-003 | 1.0558-003 | 1.0347-003 |
| 1.0540-003 | 1.0411-003 | 1.0202-003 |
| 1.0395-003 | 1.0268-003 | 1.0061-003 |
| 1.0253-003 | 1.0128-003 | 9.9230-004 |
| 1.0113-003 | 9.9909-004 | 9.7875-004 |


| $=0.6$ | $=0.8$ | 10 $\beta$ |
| :---: | :---: | :---: |
| 2.2417-003 | 2.1505-003 | 136 |
| 2.1994-003 | 2.1115-003 | 137 |
| 2.1578-003 | 2.0682-003 | 138 |
| 2.1171-003 | 2.0313-003 | 139 |
| 2.0773-003 | 1.9960-003 | 140 |
| 2.0384-003 | 1.9555-003 | 141 |
| 2.0004-003 | 1.9205-003 | 142 |
| 1.9634-003 | 1.8890-003 | 143 |
| 1.9274-003 | 1.8519-003 | 144 |
| 1.8924-003 | 1.8189-003 | 145 |
| 1.8585-003 | 1.7909-003 | 146 |
| 1.8255-003 | 1.7575-003 | 147 |
| 1.7935-003 | 1.7264-003 | 148 |
| 1.7625-003 | 1.7015-003 | 149 |
| 1.7325-003 | 1.6717-003 | 150 |
| 1.7033-003 | 1.6424-003 | 151 |
| 1.6751-003 | 1.6199-003 | 152 |
| 1.6476-003 | 1.5936-003 | 153 |
| 1.6209-003 | 1.5657-003 | 154 |
| 1.5951-003 | 1.5449-003 | 155 |
| 1.5699-003 | 1.5216-003 | 156 |
| 1.5453-003 | 1.4949-003 | 157 |
| 1.5214-003 | 1.4750-003 | 158 |
| 1.4981-003 | 1.4541-003 | 159 |
| $1.4753-003$ | 1.4284-003 | 160 |
| 1.4530-003 | 1.4089-003 | 161 |
| 1.4312-003 | 1.3903-003 | 162 |
| 1.4098-003 | 1.3660-003 | 163 |
| 1.3888-003 | 1.3467-003 | 164 |
| 1.3682-003 | 1.3301-003 | 165 |
| 1.3479-003 | 1.3072-003 | 166 |
| 1.3280-003 | 1.2877-003 | 167 |
| 1.3084-003 | 1.2722-003 | 168 |
| 1.2892-003 | 1.2506-003 | 169 |
| $1.2703-003$ | 1.2307-003 | 170 |
| 1.2518-003 | 1.2162-003 | 171 |
| 1.2336-003 | 1.1963-003 | 172 |
| 1.2158-003 | 1.1766-003 | 173 |
| 1.1983-003 | 1.1633-003 | 174 |
| 1.1812-003 | 1.1459-003 | 175 |
| 1.1643-003 | .1.1266-003 | 176 |
| 1.1478-003 | 1.1141-003 | 177 |
| 1.1315-003 | 1.0987-003 | 178 |
| 1.1155-003 | 1.0798-003 | 179 |
| $1.0997-003$ | $1.0677-003$ | 180 |
| 1.0843-003 | 1.0546-003 | 181 |
| 1.0691-003 | $1.0367-003$ | 182 |
| 1.0542-003 | 1.0250-003 | 183 |
| 1.0396-003 | 1.0137-003 | 184 |
| $1.0253-003$ | 9.9633-004 | 185 |
| 1.0114-003 | 9.8381-004 | 186 |
| 9.9779-004 | 9.7351-004 | 187 |
| 9.8441-004 | 9.5693-004 | 188 |
| 9.7131-004 | 9.4391-004 | 189 |
| 9.5849-004 | 9.3483-004 | 190 |


| $\mathrm{r}_{0} / \mathrm{D}=0$ | $=0.2$ | $=0.4$ |
| :---: | :---: | :---: |
| 9.9766-004 | 9.8567-004 | 9.6548-004 |
| 9.8425-004 | 9.7253-004 | 9.5249-004 |
| 9.7109-004 | 9.5966-004 | 9.3977-004 |
| 9.5817-004 | 9.4703-004 | 9.2732-004 |
| $9.4550-004$ | 9.3465-004 | 9.1513-004 |
| 9.3306-004 | 9.2251-004 | 9.0321-004 |
| 9.2085-004 | 9.1060-004 | 8.9154-004 |
| 9.0887-004 | 8.9892-004 | 8.8012-004 |
| 8.9710-004 | 8.8745-004 | 8.6895-004 |
| 8.8554-004 | 8.7620-004 | 8.5801-004 |
| 8.7420-004 | 8.6516-004 | 8.4729-004 |
| 8.6305-004 | 8.5433-004 | 8.3680-004 |
| 8.5211-004 | 8.4370-004 | 8.2653-004 |
| 8.4136-004 | 8.3328-004 | 8.1645-004 |
| 8.3080-004 | 8.2306-004 | 8.0657-004 |
| 8.2042-004 | 8.1304-004 | 7.9688-004 |
| 8.1023-004 | 8.0321-004 | 7.8738-004 |
| 8.0021-004 | 7.9359-004 | 7.7805-004 |
| 7.9037-004 | 7.8417-004 | 7.6890-004 |
| 7.8069-004 | 7.7495-004 | 7.5993-004 |
| 7.7118-004 | 7.6593-004 | 7.5114-004 |
| 7.6184-004 | 7.5711-004 | 7.4254-004 |
| 7.5265-004 | 7.4850-004 | 7.3413-004 |
| 7.4361-004 | 7.4008-004 | 7.2591-004 |
| 7.3473-004 | 7.3186-004 | 7.1787-004 |
| 7.2600-004 | 7.2384-004 | 7.1001-004 |
| 7.1741-004 | 7.1600-004 | 7.0233-004 |
| 7.0896-004 | 7.0834-004 | 6.9480-004 |
| 7.0066-004 | 7.0084-004 | 6.8741-004 |
| 6.9249-004 | 6.9351-004 | 6.8016-004 |
| 6.8445-004 | 6.8631-004 | 6.7301-004 |
| 6.7654-004 | 6.7925-004 | 6.6596-004 |
| 6.6876-004 | 6.7229-004 | 6.5898-004 |
| 6.6111-004 | 6.6543-004 | 6.5206-004 |
| 6.5358-004 | 6.5864-004 | 6.4518-004 |
| 6.4617-004 | 6.5191-004 | 6.3833-004 |
| 6.3887-004 | 6.4523-004 | 6.3149-004 |
| 6.3169-004 | 6.3858-004 | 6.2465-004 |
| 6.2463-004 | 6.3194-004 | 6.1781-004 |
| 6.1767-004 | 6.2530-004 | 6.1095-004 |
| 6.1082-004 | 6.1867-004 | 6.0409-004 |
| 6.0408-004 | 6.1202-004 | 5.9721-004 |
| 5.9745-004 | 6.0536-004 | 5.9033-004 |
| 5.9091-004 | 5.9870-004 | 5.8346-004 |
| $5.8448-004$ | 5.9203-004 | 5.7661-004 |
| 5.7814-004 | 5.8536-004 | 5.6981-004 |
| 5.7190-004 | 5.7871-004 | 5.6306-004 |
| 5.6575-004 | 5.7209-004 | 5.5640-004 |
| 5.5970-004 | 5.6551-004 | 5.4985-004 |
| 5.5373-004 | 5.5898-004 | 5.4343-004 |
| 5.4786-004 | 5.5253-004 | 5.3716-004 |
| 5.4207-004 | 5.4617-004 | 5.3107-004 |
| 5.3637-004 | 5.3991-004 | 5.2517-004 |
| 5.3075-004 | 5.3377-004 | 5.1948-004 |
| 5.2522-004 | 5.2775-004 | 5.1400-004 |


| $=0.6$ | $=0.8$ | $10 \beta$ |
| :---: | :---: | :---: |
| 9.4582-004 | 9.1956-004 | 191 |
| 9.3337-004 | 9.0622-004 | 192 |
| 9.2114-004 | 8.9832-004 | 193 |
| 9.0905-004 | 8.8498-004 | 194 |
| 8.9714-004 | 8.7181-004 | 195 |
| 8.8547-004 | 8.6485-004 | 196 |
| 8.7397-004 | 8.5314-004 | 197 |
| 8.6268-004 | 8.3942-004 | 198 |
| 8.5167-004 | 8.3222-004 | 199 |
| 8.4087-004 | 8.2177-004 | 200 |
| 8.3025-004 | 8.0788-004 | 201 |
| 8.1991-004 | 8.0060-004 | 202 |
| 8.0975-004 | 7.9175-004 | 203 |
| 7.9974-004 | 7.7797-004 | 204 |
| 7.8996-004 | 7.7011-004 | 205 |
| 7.8035-004 | 7.6243-004 | 206 |
| 7.7086-004 | 7.4890-004 | 207 |
| 7.6159-004 | 7.4041-004 | 208 |
| 7.5251-004 | 7.3411-004 | 209 |
| 7.4357-004 | 7.2164-004 | 210 |
| 7.3485-004 | 7.1292-004 | 211 |
| 7.2636-004 | 7.0807-004 | 212 |
| 7.1801-004 | 6.9702-004 | 213 |
| 7.0988-004 | 6.8795-004 | 214 |
| 7.0196-004 | 6.8401-004 | 215 |
| 6.9417-004 | 6.7430-004 | 216 |
| 6.8654-004 | 6.6469-004 | 217 |
| 6.7910-004 | 6.6118-004 | 218 |
| 6.7174-004 | 6.5294-004 | 219 |
| 6.6447-004 | 6.4302-004 | 220 |
| 6.5735-004 | 6.3973-004 | 221 |
| 6.5028-004 | 6.3308-004 | 222 |
| 6.4325-004 | 6.2301-004 | 223 |
| 6.3632-004 | 6.1949-004 | 224 |
| 6.2941-004 | 6.1419-004 | 225 |
| 6.2248-004 | 6.0400-004 | 226 |
| 6.1562-004 | 5.9978-004 | 227 |
| 6.0876-004 | 5.9553-004 | 228 |
| 6.0186-004 | 5.8535-004 | 229 |
| 5.9498-004 | 5.8011-004 | 230 |
| 5.8812-004 | 5.7664-004 | 231 |
| 5.8121-004 | 5.6671-004 | 232 |
| 5.7434-004 | 5.6033-004 | 233 |
| 5.6752-004 | 5.5742-004 | 234 |
| 5.6069-004 | 5.4806-004 | 235 |
| 5.5395-004 | 5.4062-004 | 236 |
| 5.4733-004 | 5.3810-004 | 237 |
| 5.4080-004 | 5.2970-004 | 238 |
| 5.3441-004 | 5.2149-004 | 239 |
| 5.2825-004 | 5.1931-004 | 240 |
| 5.2226-004 | 5.1230-004 | 241 |
| 5.1649-004 | 5.0377-004 | 242 |
| 5.1103-004 | 5.0189-004 | 243 |
| 5.0580-004 | 4.9664-004 | 244 |
| 5.0082-004 | 4.8824-004 | 245 |

LOW FREQUENCY COMPONENT OF THE ELECTRIC MICROFIELDS

| $\mathrm{r}_{0} / \mathrm{D}=0$ | $=0.2$ | $=0.4$ |
| :---: | :---: | :---: |
| 5.1976-004 | 5.2185-004 | 5.0873-004 |
| 5.1438-004 | 5.1609-004 | 5.0366-004 |
| 5.0909-004 | 5.1046-004 | 4.9878-004 |
| 5.0387-004 | 5.0495-004 | 4.9408-004 |
| 4.9872-004 | 4.9956-004 | 4.8952-004 |
| 4.9364-004 | 4.9364-004 | 4.9364-004 |
| 4.8864-004 | 4.8864-004 | 4.8864-004 |
| 4.8371-004 | 4.8371-004 | 4.8371-004 |
| 4.7885-004 | 4.7885-004 | 4.7885-004 |
| 4.7406-004 | 4.7406-004 | 4.7406-004 |
| 4.6934-004 | 4.6934-004 | 4.6934-004 |
| 4.6468-004 | 4.6468-004 | 4.6468-004 |
| 4.6008-004 | 4.6008-004 | 4.6008-004 |
| 4.5555-004 | 4.5555-004 | 4.5555-004 |
| 4.5108-004 | 4.5108-004 | 4.5108-004 |
| 4.4667-004 | 4.4667-004 | 4.4667-004 |
| 4.4233-004 | 4.4233-004 | 4.4233-004 |
| 4.3804-004 | 4.3804-004 | 4.3804-004 |
| 4.3381-004 | 4.3381-004 | 4.3381-004 |
| 4.2964-004 | 4.2964-004 | 4.2964-004 |
| 4.2552-004 | 4.2552-004 | 4.2552-004 |
| 4.2146-004 | 4.2146-004 | 4.2146-004 |
| 4.1745-004 | 4.1745-004 | 4.1745-004 |
| 4.1350-004 | 4.1350-004 | 4.1350-004 |
| 4.0960-004 | 4.0960-004 | 4.0960-004 |
| 4.0575-004 | 4.0575-004 | 4.0575-004 |
| 4.0195-004 | 4.0195-004 | 4.0195-004 |
| 3.9820-004 | 3.9820-004 | 3.9820-004 |
| 3.9450-004 | 3.9450-004 | 3.9450-004 |
| 3.9084-004 | 3.9084-004 | 3.9084-004 |
| 3.8724-004 | 3.8724-004 | 3.8724-004 |
| 3.8368-004 | 3.8368-004 | 3.8368-004 |
| 3.8017-004 | 3.8017-004 | 3.8017-004 |
| 3.7670-004 | 3.7670-004 | 3.7670-004 |
| $3.7327-004$ | 3.7327-004 | 3.7327-004 |
| 3.6989-004 | 3.6989-004 | 3.6989-004 |
| 3.6656-004 | 3.6656-004 | 3.6656-004 |
| 3.6326-004 | 3.6326-004 | 3.6326-004 |
| 3.6001-004 | 3.6001-004 | 3.6001-004 |
| 3.5679-004 | 3.5679-004 | 3.5679-004 |
| 3.5362-004 | 3.5362-004 | 3.5362-004 |
| 3.5049-004 | 3.5049-004 | 3.5049-004 |
| 3.4739-004 | 3.4739-004 | 3.4739-004 |
| 3.4434-004 | 3.4434-004 | 3.4434-004 |
| 3.4132-004 | 3.4132-004 | 3.4132-004 |
| 3.3834-004 | 3.3834-004 | 3.3834-004 |
| 3.3539-004 | 3.3539-004 | 3.3539-004 |
| 3.3248-004 | 3.3248-004 | 3.3248-004 |
| 3.2961-004 | 3.2961-004 | 3.2961-004 |
| 3.2677-004 | 3.2677-004 | 3.2677-004 |
| 3.2396-004 | 3.2396-004 | 3.2396-004 |
| 3.2119-004 | 3.2119-004 | 3.2119-004 |
| 3.1845-004 | 3.1845-004 | 3.1845-004 |
| 3.1575-004 | 3.1575-004 | 3.1575-004 |
| 3.1307-004 | 3.1307-004 | 3.1307-004 |

$=0.6$
$4.9617-004$
$4.9176-004$
$4.8754-004$
$4.8358-004$
$4.7979-004$
$4.7607-004$
$4.7247-004$
$4.6890-004$
$4.6527-004$
$4.6160-004$
$4.5786-004$
$4.5395-004$
$4.4992-004$
$4.4579-004$
$4.4148-004$
$4.3704-004$
$4.3254-004$
$4.2793-004$
$4.2324-004$
$4.1858-004$
$4.1390-004$
$4.0923-004$
$4.0467-004$
$4.0019-004$
$3.9577-004$
$3.9153-004$
$3.8743-004$
$3.8342-004$
$3.7960-004$
$3.7594-004$
$3.7238-004$
$3.6898-004$
$3.6573-004$
$3.6255-004$
$3.5950-004$
$3.5657-004$
$3.5369-004$
$3.5087-004$
$3.4817-004$
$3.4547-004$
$3.4279-004$
$3.4019-004$
$3.3757-004$
$3.3493-004$
$3.3234-004$
$3.2972-004$
$3.2704-004$
$3.2442-004$
$3.2178-004$
$3.1910-004$
$3.1647-004$
$3.1387-004$
$3.1124-004$
$3.0867-004$
$3.0614-004$

| $=0.8$ | $10 \beta$ |
| :---: | :---: |
| 4.8657-004 | 246 |
| 4.8319-004 | 247 |
| 4.7522-004 | 248 |
| 4.7346-004 | 249 |
| 4.7171-004 | 250 |
| 4.6423-004 | 251 |
| 4.6187-004 | 252 |
| 4.6119-004 | 253 |
| 4.5408-004 | 254 |
| 4.5058-004 | 255 |
| 4.5034-004 | 256 |
| 4.4350-004 | 257 |
| 4.3857-004 | 258 |
| 4.3829-004 | 259 |
| 4.3188-004 | 260 |
| 4.2559-004 | 261 |
| 4.2505-004 | 262 |
| 4.1938-004 | 263 |
| 4.1209-004 | 264 |
| 4.1122-004 | 265 |
| 4.0666-004 | 266 |
| 3.9887-004 | 267 |
| 3.9766-004 | 268 |
| 3.9446-004 | 269 |
| 3.8665-004 | 270 |
| 3.8501-004 | 271 |
| 3.8323-004 | 272 |
| 3.7578-004 | 273 |
| 3.7359-004 | 274 |
| 3.7307-004 | 275 |
| 3.6628-004 | 276 |
| 3.6339-004 | 277 |
| 3.6384-004 | 278 |
| 3.5789-004 | 279 |
| 3.5415-004 | 280 |
| 3.5515-004 | 281 |
| 3.5009-004 | 282 |
| 3.4539-004 | 283 |
| 3.4644-004 | 284 |
| $3.4231-004$ | 285 |
| 3.3669-004 | 286 |
| 3.3746-004 | 287 |
| 3.3441-004 | 288 |
| 3.2818-004 | 289 |
| 3.2845-004 | 290 |
| 3.2648-004 | 291 |
| 3.1980-004 | 292 |
| 3.1914-004 | 293 |
| 3.1799-004 | 294 |
| 3.1103-004 | 295 |
| 3.0924-004 | 296 |
| 3.0888-004 | 297 |
| 3.0222-004 | 298 |
| 2.9956-004 | 299 |
| 3.0015-004 | 300 |

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    ${ }^{*}$ A Also at the Joint Institute for Laboratory Astrophysics and Department of Physics and Astrophysics, University of Colorado. Boulder. Colo. 80302.
    ${ }^{1}$ Figures in brackets refer to the literature references on p. 14.

